

Classical Mechanics and Statistical/Thermodynamics

August 2008

Possibly Useful Information

Handy Integrals:

$$\int_0^{\infty} x^n e^{-\alpha x} dx = \frac{n!}{\alpha^{n+1}}$$

$$\int_0^{\infty} e^{-\alpha x^2} dx = \frac{1}{2} \sqrt{\frac{\pi}{\alpha}}$$

$$\int_0^{\infty} x e^{-\alpha x^2} dx = \frac{1}{2\alpha}$$

$$\int_0^{\infty} x^2 e^{-\alpha x^2} dx = \frac{1}{4} \sqrt{\frac{\pi}{\alpha^3}}$$

$$\int_{-\infty}^{\infty} e^{i a x - b x^2} dx = \sqrt{\frac{\pi}{b}} e^{-a^2/4b}$$

Geometric Series:

$$\sum_{n=0}^{\infty} x^n = \frac{1}{1-x} \quad \text{for } |x| < 1$$

Stirling's approximation:

$$n! \approx \left(\frac{n}{e}\right)^n \sqrt{2\pi n}$$

Riemann and related functions:

$$\sum_{n=1}^{\infty} \frac{1}{n^p} \equiv \zeta(p)$$

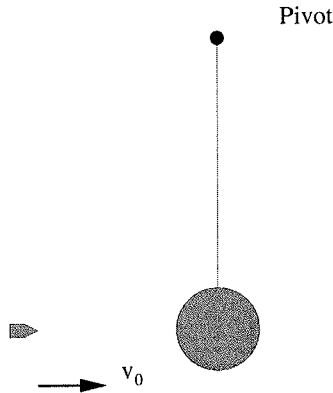
$$\sum_{n=1}^{\infty} \frac{z^p}{n^p} \equiv g_p(z) \quad \sum_{n=1}^{\infty} (-1)^p \frac{z^p}{n^p} \equiv f_p(z)$$

$$g_p(1) = \zeta(p) \quad f_p(1) = \zeta(-p)$$

$$\begin{array}{ll} \zeta(1) = \infty & \zeta(-1) = 0.0833333 \\ \zeta(2) = 1.64493 & \zeta(-2) = 0 \\ \zeta(3) = 1.20206 & \zeta(-3) = 0.0083333 \\ \zeta(4) = 1.08232 & \zeta(-4) = 0 \end{array}$$

Classical Mechanics

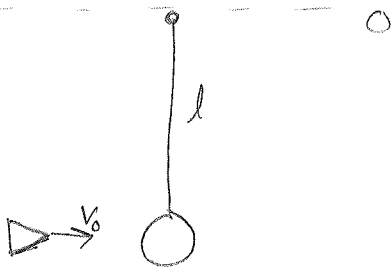
1. **The ballistic pendulum:** Consider a pendulum with a bob of mass m connected to a frictionless pivot by an ideal massless rigid rod of length ℓ . A projectile of mass ϵm ($0 < \epsilon \ll 1$) moving horizontally at speed v_0 hits the center of the bob, as shown. When it strikes, it becomes imbedded in the bob.



- (a) What is the minimum initial speed of the projectile such that the pendulum will make a full rotation? (2 points)
- (b) The rod is replaced by an ideal massless non-rigid string. What is the minimum initial speed of the projectile such that the pendulum will make a full revolution without the string going slack? (3 points)
- (c) Now assume that projectile rebounds elastically from the bob in the horizontal direction. What is the minimum initial speed of the projectile such that the pendulum will make a full revolution without the string going slack? (2 points)
- (d) Finally, assume that the projectile passes completely through the pendulum bob, in a time $t \ll \sqrt{\ell/g}$. After it exits, it carries with it some of the original mass of the bob, such that the exiting projectile now has a mass $2\epsilon m$ and moves at a speed $3v_0/4$. What is the minimum initial speed of the projectile such that the pendulum will make a full revolution without the string going slack? (3 points)

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Classical #1



$$m_b = 6m$$

$$m_p = m$$

a) * Collision (Inelastic)

$$\cancel{\frac{1}{2} m v_0^2 = \frac{1}{2} m (1+e) v^2}$$

$$\cancel{\frac{e}{1+e} v_0^2 = v^2}$$

$$e m v_0 = (1+e) m v$$

$$\frac{e}{1+e} v_0 = v$$

* Conservation of energy

$$\frac{1}{2} m (1+e) \left(\frac{e}{1+e} v_0 \right)^2 - m g l = m g l + \frac{1}{2} m (1+e) v^2$$

$$\frac{1}{2} m \frac{e^2}{1+e} v_0^2 = 2 m g l + \frac{1}{2} m (1+e) v^2$$

$$\frac{e^2}{1+e} v_0^2 = 4 m g l$$

2. The isotropic harmonic oscillator.

- (a) Write the Lagrangian for a point mass m moving under the influence of an isotropic 3-dimensional harmonic oscillator potential

$$V(x, y, z) = \frac{k}{2}(x^2 + y^2 + z^2).$$

There is no external gravitational field. (1 point)

- (b) Using the Lagrange equations of motion show that angular momentum is conserved. i.e.,

$$\frac{d}{dt}\mathbf{L} = \frac{d}{dt}(\mathbf{r} \times m\mathbf{v}) = 0.$$

Because the Lagrangian is invariant under rotations about the origin, you can choose coordinates so that motion is constrained to the x-y plane, i.e., the angular momentum points in the z direction. (3 points)

- (c) For 2-dimensional motion in the x-y plane choose cylindrical polar coordinates and proceed to solve the Lagrange equations of motion. You can leave the solution for $r(t)$ as an integral of the form $t = \int f(r)dr$. (Don't forget to use conservation of energy, E_0 .) (3 points)
- (d) Compute the minimum and maximum values or the radial coordinate r as functions of the constants m, E_0, k, L^z . (3 points)

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Classical # 2

a) Write Lagrangian for a point mass under the influence of the following potential:

$$V(x, y, z) = \frac{k}{2}(x^2 + y^2 + z^2) \quad * \text{ No external gravity}$$

$$\mathcal{L} = T - V$$

$$= \frac{1}{2} m v^2 - \frac{1}{2} k (x^2 + y^2 + z^2)$$

$$= \frac{1}{2} m (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) - \frac{1}{2} k (x^2 + y^2 + z^2)$$

b) Use Lagrange equations of motion to show angular momentum conservation

i.e. $\frac{d}{dt} \vec{L} = \frac{d}{dt} (\vec{r} \times m \vec{v}) = 0$

$$\frac{\partial \mathcal{L}}{\partial x} = \frac{\partial}{\partial t} \frac{\partial \mathcal{L}}{\partial \dot{x}}$$

$$\frac{\partial \mathcal{L}}{\partial y} = \frac{\partial}{\partial t} \frac{\partial \mathcal{L}}{\partial \dot{y}}$$

$$\frac{\partial \mathcal{L}}{\partial z} = \frac{\partial}{\partial t} \frac{\partial \mathcal{L}}{\partial \dot{z}}$$

$$\Rightarrow -kx = \frac{\partial}{\partial t} (m\dot{x})$$

$$\Rightarrow -ky = \frac{\partial}{\partial t} (m\dot{y})$$

$$\Rightarrow -kz = \frac{\partial}{\partial t} (m\dot{z})$$

$$-kx = m\ddot{x}$$

$$-ky = m\ddot{y}$$

$$-kz = m\ddot{z}$$

$$\vec{r} \times m \vec{v} = \begin{vmatrix} \hat{x} & \hat{y} & \hat{z} \\ x & y & z \\ m\dot{x} & m\dot{y} & m\dot{z} \end{vmatrix} = \langle ym\dot{z} - zm\dot{y}, xm\dot{z} - zm\dot{x}, xmy - ymx \rangle$$

$$= m \langle y\dot{z} - z\dot{y}, x\dot{z} - z\dot{x}, x\dot{y} - y\dot{x} \rangle$$

$$\frac{d}{dt} (\vec{r} \times m \vec{v}) = m \langle \dot{y}\dot{z} + y\ddot{z} - (\dot{z}\dot{y} + z\ddot{y}), \dot{x}\dot{z} + x\ddot{z} - (\dot{z}\dot{x} + z\ddot{x}), \dot{x}\dot{y} + x\ddot{y} - (\dot{y}\dot{x} + y\ddot{x}) \rangle$$

$$= m \langle y\ddot{z} - z\ddot{y}, x\ddot{z} - z\ddot{x}, x\ddot{y} - y\ddot{x} \rangle$$

* substituting from above

$$= m \left(\frac{-k}{m} \right) \langle yz - zy, xz - zx, xy - yx \rangle$$

$$= -k \langle 0, 0, 0 \rangle$$

$$= 0$$

$$\therefore \frac{d}{dt} (\vec{r} \times m \vec{v}) = \frac{d}{dt} \vec{L} = 0 \checkmark$$

c) For 2-D motion in x-y plane, use cylindrical polar coordinates to solve Lagrange eqns of motion. Leave $\dot{r}(t)$ as an integral of form $t = \int f(r) dr$

* Hint: use conservation of energy E_0

- In cylindrical polar: $x = r \cos \phi$ $\dot{x} = \dot{r} \cos \phi - r \sin \phi \dot{\phi}$
 $y = r \sin \phi$ $\dot{y} = \dot{r} \sin \phi + r \cos \phi \dot{\phi}$

$$\Rightarrow \mathcal{L} = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}^2) - \frac{1}{2} k r^2 \quad (2-D, z=0)$$

$$E_0 = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}^2) + \frac{1}{2} k r^2$$

$$\frac{\partial \mathcal{L}}{\partial r} - \frac{\partial}{\partial t} \frac{\partial \mathcal{L}}{\partial \dot{r}} = 0$$

$$\frac{\partial \mathcal{L}}{\partial \phi} - \frac{\partial}{\partial t} \left(\frac{\partial \mathcal{L}}{\partial \dot{\phi}} \right) = 0$$

$$0 = kr + m r \dot{\phi}^2 - m \ddot{r}$$

$$0 = -2m r \dot{r} \dot{\phi} - m r^2 \ddot{\phi}$$

* Solving E_0 for $\dot{\phi}$ yields

$$E_0 = \frac{1}{2} m \dot{r}^2 + \frac{1}{2} m r^2 \dot{\phi}^2 + \frac{1}{2} k r^2$$

$$\frac{2E_0}{m} = \dot{r}^2 + r^2 \dot{\phi}^2 + \frac{k}{m} r^2$$

$$\frac{2E_0}{m} - \frac{k}{m} r^2 - \dot{r}^2 = r^2 \dot{\phi}^2$$

$$\frac{2E_0}{m r^2} - \frac{k}{m} - \frac{\dot{r}^2}{r^2} = \dot{\phi}^2$$

$$\Rightarrow 0 = kr + m r \left(\frac{2E_0}{m r^2} - \frac{k}{m} - \frac{\dot{r}^2}{r^2} \right) + \frac{1}{2} k r^2$$

$$= kr + \frac{2E_0}{r} - kr - \frac{m \dot{r}^2}{r} + \frac{1}{2} k r^2$$

$$0 = \frac{2E_0}{r} - \frac{m \dot{r}^2}{r} + \frac{1}{2} k r^2$$

$$\frac{m \dot{r}^2}{r} = \frac{2E_0}{r} + \frac{1}{2} k r^2$$

$$\dot{r}^2 = \frac{2E_0}{m} + \frac{1}{2} \frac{k}{m} r^3$$

$$\Rightarrow t = \int \left(\frac{2E_0}{m} + \frac{1}{2} \frac{k}{m} r^3 \right)^{1/2} dr$$

3. Consider a particle attracted by a fixed gravitating body while also in a uniform gravitational field oriented along the z -axis. The potential energy is of the form:

$$V(r, z) = -m \left(\frac{k}{r} + g z \right)$$

where m is the particle's mass, k and g are constants, and r is the standard radial coordinate:

$$r \equiv \sqrt{x^2 + y^2 + z^2}$$

You are to examine the problem in *cylindrical parabolic coordinates* defined by

$$\begin{aligned}\zeta &\equiv r + z \\ \eta &\equiv r - z \\ \phi &\equiv \arctan y/x\end{aligned}$$

In these coordinates we may write the Cartesian coordinates as:

$$\begin{aligned}x &= \sqrt{\zeta\eta} \cos \phi \\ y &= \sqrt{\zeta\eta} \sin \phi \\ z &= \frac{1}{2}(\zeta - \eta)\end{aligned}$$

- (a) Show that the kinetic energy, T , is given by:

$$T = \frac{m}{8} \left[\left(1 + \frac{\zeta}{\eta} \right) \dot{\eta}^2 + \left(1 + \frac{\eta}{\zeta} \right) \dot{\zeta}^2 \right] + \frac{m}{2} \zeta \eta \dot{\phi}^2$$

in these coordinates. (2 points)

- (b) What are the canonical momenta, p_ζ , p_η , and p_ϕ , expressed in cylindrical parabolic coordinates? (2 points)
- (c) Use Hamilton-Jacobi theory to find the constants of the motion. *Hint:* While the total energy E does not separate in these coordinates, $E(\zeta + \eta)$ can be used to produce a quantity that **does** separate. (3 points)
- (d) What is Hamilton's characteristic function associated with ϕ ? (1 point)
- (e) Express Hamilton's characteristic functions associated with ζ , η as definite integrals. (2 points)

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Classical #3

For a particle attracted by a fixed gravitating body while also in a uniform gravitational field oriented along the z -axis [$V(r,z) = -m(\frac{k}{r} + gz)$] where $r = \sqrt{x^2 + y^2 + z^2}$ and using cylindrical parabolic coordinates:

$$\begin{aligned} \rho &= r + z & x &= \sqrt{\rho\eta} \cos \phi \\ \eta &= r - z & y &= \sqrt{\rho\eta} \sin \phi \\ \phi &= \arctan(y/x) & z &= \frac{1}{2}(\rho - \eta) \end{aligned}$$

a) Show that the kinetic energy, T , is: $T = \frac{m}{8} \left[\left(1 + \frac{\rho}{\eta}\right) \dot{\eta}^2 + \left(1 + \frac{\eta}{\rho}\right) \dot{\rho}^2 \right] + \frac{m}{2} \rho \eta \dot{\phi}^2$

$$\begin{aligned} T &= \frac{1}{2} m \dot{\mathbf{v}}^2 \\ &= \frac{m}{2} (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) \end{aligned}$$

* From above,

$$\dot{x} = \frac{1}{2} \rho^{-1/2} \eta^{1/2} \cos \phi \dot{\rho} + \frac{1}{2} \rho^{1/2} \eta^{-1/2} \cos \phi \dot{\eta} - \rho^{1/2} \eta^{1/2} \sin \phi \dot{\phi}$$

$$\begin{aligned} \dot{x}^2 &= \frac{1}{4} \rho^{-1} \eta \cos^2 \phi \dot{\rho}^2 + \frac{1}{4} \rho \eta^{-1} \cos^2 \phi \dot{\eta}^2 + \rho \eta \sin^2 \phi \dot{\phi}^2 + \frac{1}{2} \cos^2 \phi \dot{\rho} \dot{\eta} - \rho \cos \phi \sin \phi \dot{\rho} \dot{\phi} \\ &\quad - \rho \cos \phi \sin \phi \dot{\eta} \dot{\phi} \end{aligned}$$

$$\dot{y} = \frac{1}{2} \rho^{-1/2} \eta^{1/2} \sin \phi \dot{\rho} + \frac{1}{2} \rho^{1/2} \eta^{-1/2} \sin \phi \dot{\eta} + \rho^{1/2} \eta^{1/2} \cos \phi \dot{\phi}$$

$$\begin{aligned} \dot{y}^2 &= \frac{1}{4} \rho^{-1} \eta \sin^2 \phi \dot{\rho}^2 + \frac{1}{4} \rho \eta^{-1} \sin^2 \phi \dot{\eta}^2 + \rho \eta \cos^2 \phi \dot{\phi}^2 + \frac{1}{2} \sin^2 \phi \dot{\rho} \dot{\eta} + \rho \sin \phi \cos \phi \dot{\rho} \dot{\phi} \\ &\quad + \rho \sin \phi \cos \phi \dot{\eta} \dot{\phi} \end{aligned}$$

$$\dot{z} = \frac{1}{2} (\dot{\rho} - \dot{\eta})$$

$$\dot{z}^2 = \frac{1}{4} (\dot{\rho}^2 - 2\dot{\rho}\dot{\eta} + \dot{\eta}^2)$$

$$\Rightarrow \dot{x}^2 + \dot{y}^2 + \dot{z}^2 = \frac{1}{4} \rho^{-1} \eta \dot{\rho}^2 + \frac{1}{4} \rho \eta^{-1} \dot{\eta}^2 - \rho \eta \sin^2 \phi \dot{\phi}^2 + \rho \eta \cos^2 \phi \dot{\phi}^2 + \frac{1}{2} \dot{\rho} \dot{\eta} + \frac{1}{4} \dot{\rho}^2 - \frac{1}{2} \dot{\rho} \dot{\eta} + \frac{1}{4} \dot{\eta}^2$$

$$= \frac{1}{4} \frac{\rho}{\eta} \dot{\rho}^2 + \frac{1}{4} \frac{\rho}{\eta} \dot{\eta}^2 - \rho \eta \sin^2 \phi \dot{\phi}^2 + \rho \eta \cos^2 \phi \dot{\phi}^2 + \frac{1}{4} \dot{\rho}^2 + \frac{1}{4} \dot{\eta}^2$$

$$= \frac{1}{4} \left[\dot{\rho}^2 \left(\frac{\rho}{\eta} + 1 \right) + \dot{\eta}^2 \left(\frac{\rho}{\eta} + 1 \right) \right] + \rho \eta \dot{\phi}^2 (\cos^2 \phi + \sin^2 \phi)$$

$$\Rightarrow T = \frac{1}{2} m (\dot{x}^2 + \dot{y}^2 + \dot{z}^2)$$

$$= \frac{m}{8} \left[\left(1 + \frac{\rho}{\eta}\right) \dot{\eta}^2 + \left(1 + \frac{\eta}{\rho}\right) \dot{\rho}^2 \right] + \frac{m}{2} \rho \eta \dot{\phi}^2$$

b) Find the canonical momenta in cylindrical parabolic coordinates

$$P_q = \frac{\partial \mathcal{L}}{\partial \dot{q}}$$

$$\Rightarrow \mathcal{L} = T - V$$

$$V = -m\left(\frac{k}{r} + gz\right), \quad r = \sqrt{x^2 + y^2 + z^2}$$

$$\Rightarrow r = \sqrt{zn \cos^2 \varphi + zn \sin^2 \varphi + \frac{1}{4}(z-n)^2}$$

$$= \sqrt{zn + \frac{1}{4}z^2 - \frac{1}{2}zn + \frac{1}{4}n^2}$$

$$= \sqrt{\frac{1}{4}z^2 + \frac{1}{2}zn + \frac{1}{4}n^2}$$

$$= \sqrt{\frac{1}{4}(z^2 + 2zn + n^2)}$$

$$= \frac{1}{2}(z+n)$$

$$\Rightarrow V = -m\left(\frac{2k}{z+n} + g\left(\frac{z+n}{2}\right)\right)$$

$$\Rightarrow \mathcal{L} = T - V$$

$$= \frac{m}{8} \left[\left(1 + \frac{z}{n}\right) \dot{n}^2 + \left(1 + \frac{n}{z}\right) \dot{z}^2 \right] + \frac{m}{2} zn \dot{\varphi}^2 + m \left[\frac{2k}{z+n} + \frac{g(z+n)}{2} \right]$$

$$P_\varphi = \frac{\partial \mathcal{L}}{\partial \dot{\varphi}}$$

$$= mzn\dot{\varphi}$$

$$P_n = \frac{\partial \mathcal{L}}{\partial \dot{n}}$$

$$= \frac{m}{4} \left(1 + \frac{z}{n}\right) \dot{n}$$

$$P_z = \frac{\partial \mathcal{L}}{\partial \dot{z}}$$

$$= \frac{m}{4} \left(1 + \frac{n}{z}\right) \dot{z}$$

$$c) \quad H = \sum_i \dot{p}_i \dot{q}_i - \mathcal{L}$$

$$H = p_\varphi \dot{\varphi} + p_z \dot{z} + p_n \dot{n} - \mathcal{L}$$

$$= \frac{m}{2} 3n \dot{\varphi}^2 + \frac{m}{8} (1 + \frac{2}{3}) \dot{z}^2 + \frac{m}{8} (1 + \frac{3}{2}) \dot{n}^2 - m \left[\frac{2k}{3+n} + g \frac{(3+n)}{2} \right]$$

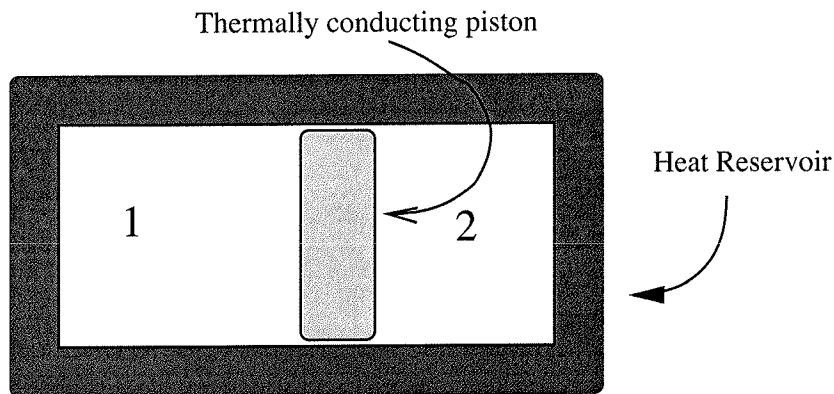
Statistical Mechanics

4. **Helmholtz Free Energy:** The Helmholtz free energy of an ideal monoatomic gas can be written as

$$F(T, V, N) = NkT \left\{ A - \log \left[T^{3/2} \frac{V}{N} \right] \right\}$$

where N is the total number of gas atoms, V is the volume, T is temperature, k is Boltzmann's constant and A is a dimensionless constant.

Consider a piston separating a system into two parts, with equal numbers of particles on the left and the right hand side. The whole system is in good thermal contact with a reservoir at constant temperature T . Initially, $V_1 = 2V_2$. The total volume, $V_{\text{tot}} = V_1 + V_2$, is fixed for this whole problem.



- (a) Calculate the equilibrium position of the piston, once it is released. You must prove your answer, and not simply assert it. (3 points)
- (b) Calculate the maximum available work the system can perform as it changes from the initial condition to the equilibrium position. (3 points)
- (c) Calculate the change in the internal energy, U of gas 1 and gas 2 in the process. (2 points)
- (d) Given your answers above, explain the source of energy for the work done during the expansion. (2 points)

5. Consider a gas of N non-interacting **one dimensional** diatomic molecules enclosed in a box of “volume” L (actually, just a length) at temperature T .

- (a) The classical energy for a single molecule is:

$$E(p_1, p_2, x_1, x_2) = \frac{p_1^2}{2m} + \frac{p_2^2}{2m} + \frac{1}{2}K(x_1 - x_2)^2$$

where p_1 and p_2 are the classical momenta of the atoms in one diatomic molecule, x_1 and x_2 are their classical positions, and K is the spring constant. Calculate the specific heat for the gas. (You should assume that $KL^2/2 \gg k_B T$, where k_B is Boltzmann’s constant.) (4 points).

- (b) In the quantum limit the energy levels of the molecule are discrete. In a semiclassical approach we can write the energy of one molecule as:

$$E(P, n) = \frac{P^2}{4m} + \hbar\omega(n + \frac{1}{2})$$

where P is the momentum of the diatomic molecule (of mass $2m$), and ω is the natural frequency of the oscillator, and n is a non-negative integer ($n \geq 0$). Calculate the specific heat. (4 points).

- (c) Calculate the high and low temperature limits of your result in (b), and explain how they relate to the result of (a). (2 points)

6. Fermions:

- (a) Show that for any non-interacting spin 1/2 fermionic system with chemical potential μ , the probability of occupying a single particle state with energy $\mu + \delta$ is the same as finding a state vacant at an energy $\mu - \delta$. (2 points)
- (b) Consider non-interacting fermions that come in two types of energy states:

$$E_{\pm}(\vec{k}) = \pm \sqrt{m^2 c^4 + \hbar^2 k^2 c^2}$$

At zero temperature all the states with negative energy (all states with energy $E_{-}(\vec{k})$) are occupied¹ and all positive energy states are empty, and that $\mu(T = 0) = 0$. Show that the result of part (a) above means that the chemical potential must remain at zero for all temperatures if particle number is to be conserved. (2 points)

- (c) Using the results of (a) and (b) above, show that the average excitation energy, the change in the energy of the system from it's energy at $T = 0$ in three dimensions is given by:

$$\Delta E \equiv E(T) - E(0) = 4V \int \frac{d\vec{k}}{(2\pi)^3} E_{+}(\vec{k}) \frac{1}{1 + e^{\beta E_{+}(\vec{k})}}$$

(2 points)

- (d) Evaluate the integral above for massless ($m = 0$) particles. (2 points)
- (e) Calculate the heat capacity of such particles. (2 points)

¹Technically this means the total energy of the system diverges. If this bothers you, you can assume some large cut-off to the wavevectors, $\hbar k_{\max} c \gg kT$, which will have no effect on your final answers.