

PHYS 6213: Advanced Particle Physics, Spring 2022

## Lecture 23, Apr 11, 2022 (Monday)

- Reading:
  - (a) Chap 12 in Collider Physics
  - (b) Chap 18 and Chap 19 in Quantum Field Theory
- Assignments:
  - (a) Problem Set 4 due Apr 12 (Tue)

## Topics for Today:

Chapter 12 Dimensional Regularization and Renormalization

12.1 Loop Integral in N Dimensions

12.2 Dimensional Regularization

12.3 Mass Renormalization in  $\lambda\phi^4$

## Topics for Next Lecture:

12.4 Coupling Renormalization in  $\lambda\phi^4$

## 12.1 Loop Integral in N Dimensions

Let's consider an integral in  $N$  dimensions<sup>a</sup>

$$I_N = \int d^N \ell F(\ell^2)$$
$$d^N \ell = |\ell|^{N-1} d|\ell| d\phi \sin \theta_1 d\theta_1 \sin^2 \theta_2 d\theta_2 \dots \sin^{N-2} \theta_{N-2} d\theta_{N-2}, \quad (1)$$

with the following values for the integration variables

$$0 \leq |\ell| \leq \infty,$$
$$0 \leq \phi \leq 2\pi,$$
$$0 \leq \theta_i \leq \pi, \quad i = 1, \dots, N - 2. \quad (2)$$

Applying the well known formula

$$\int_0^{\pi/2} (\sin \theta)^{2m-1} (\cos \theta)^{2n-1} = \frac{1}{2} \frac{\Gamma(m)\Gamma(n)}{\Gamma(m+n)} = \frac{1}{2} B(m, n), \quad (3)$$

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<sup>a</sup>*Field Theory: A Modern Primer* by Pierre Ramond

for  $\text{Re } \alpha, \text{Re } \beta > 0$ , and  $\Gamma(1/2) = \sqrt{\pi}$ , we obtain

$$I_N = \frac{\pi^{N/2}}{\Gamma(N/2)} \int_0^\infty dx x^{(N-2)/2} F(x), \quad (4)$$

where  $x = |\ell|^2 = \ell^2$ .

In general,  $F(x)$  will be of the form

$$F(x) = (x + M^2)^{-A}, \quad A = 1, 2, \dots$$

that leads to

$$\begin{aligned} & \int_0^\infty dx x^{(N-2)/2} (x^2 + M^2)^{-A} \\ &= (M^2)^{-A+N/2} \int_0^\infty dy y^{(N-2)/2} (1+y)^{-A} \\ &= (M^2)^{-A+N/2} B(N/2, A - N/2) = (M^2)^{-A+N/2} \frac{\Gamma(N/2)\Gamma(A - N/2)}{\Gamma(A)} \end{aligned}$$

which is valid for  $\text{Re}(N/2) > 0$  and  $\text{Re}(A - N/2) > 0$ .

Then, we have

$$\int \frac{d^N \ell}{(\ell^2 + M^2)^A} = \pi^{N/2} \frac{\Gamma(A - N/2)}{\Gamma(A)} \times \frac{1}{(M^2)^{A-N/2}} . \quad (5)$$

We have derived this expression for  $N$  integer,  $\text{Re}(A - N/2) > 0$ , and  $\text{Re}(N/2) > 0$ . Now we can generalize it for non-integer  $N$  by analytic continuation.

Now, by letting  $q = \ell + p$  we can write Eq. (5) in the following form,

$$\int \frac{d^N \ell}{(\ell^2 + 2\ell \cdot p + M^2)^A} = \pi^{N/2} \frac{\Gamma(A - N/2)}{\Gamma(A)} \frac{1}{(M^2 - p^2)^{A-N/2}} .$$

Next by successive differentiation of previous equation with respect to  $p_\mu$ , it is easy to obtain the formula

$$\int d^N \ell \frac{\ell_\mu}{(\ell^2 + 2\ell \cdot p + M^2)^A} = \pi^{N/2} \frac{\Gamma(A - N/2)}{\Gamma(A)} \frac{(-2p_\mu)}{(M^2 - p^2)^{A-N/2}}$$

and

$$\begin{aligned} & \int d^N \ell \frac{\ell_\mu \ell_\nu}{(\ell^2 + 2\ell \cdot p + M^2)^A} \\ &= \frac{\pi^{N/2}}{\Gamma(A)(M^2 - p^2)^{A-N/2}} \\ & \times \left[ \Gamma(A - N/2) p_\mu p_\nu + \frac{1}{2} \delta_{\mu\nu} \Gamma(A - 1 - N/2) (M^2 - p^2) \right]. \end{aligned}$$

For  $N = 4$ , the integral  $I_N$  becomes

$$I = \pi^2 \int_0^\infty d\ell^2 \ell^2 F(\ell^2). \quad (6)$$

## Representations of Beta Function

Here are some useful formulas for the beta function:

$$\begin{aligned} B(m, n) &= \frac{\Gamma(m)\Gamma(n)}{\Gamma(m+n)} \\ &= 2 \int_0^{\pi/2} (\sin \theta)^{2m-1} (\cos \theta)^{2n-1} d\theta \\ &= \int_0^1 x^{m-1} (1-x)^{n-1} dx \\ &= \int_0^\infty u^{m-1} (1+u)^{-(m+n)} du . \end{aligned}$$

## The Gamma Functions

The following formulae are very useful for dimensional regularization.

$$\begin{aligned}\Gamma(-n + \epsilon) &= \frac{(-1)^n}{n!} \left[ \frac{1}{\epsilon} + \psi_1(n + 1) + O(\epsilon) \right] \\ \Gamma(\epsilon) &= \frac{1}{\epsilon} - \gamma + O(\epsilon) \\ \Gamma(-1 + \epsilon) &= -\left[ \frac{1}{\epsilon} + 1 - \gamma + O(\epsilon) \right] \\ \Gamma(-2 + \epsilon) &= \frac{1}{2} \left[ \frac{1}{\epsilon} + 1 + \frac{1}{2} - \gamma + O(\epsilon) \right]\end{aligned}\tag{7}$$

where  $\gamma$  is the Euler constant, and

$$\begin{aligned}\psi_1(z) &= \frac{d[\ln \Gamma(z)]}{dz} = \frac{\Gamma'(z)}{\Gamma(z)} \\ \psi_1(n + 1) &= 1 + \frac{1}{2} + \dots + \frac{1}{n} - \gamma\end{aligned}\tag{8}$$



## Feynman Parametrization for Loop Integrals

$$\frac{1}{d_1 d_2} = \Gamma(2) \int_0^1 dx \frac{1}{[x(d_2 - d_1) + d_1]^2}$$

$$\frac{1}{d_1 d_2 d_3} = \Gamma(3) \int_0^1 dx \int_0^x dy \frac{1}{[x(d_2 - d_1) + y(d_3 - d_2) + d_1]^3}$$

$$= \Gamma(3) \int_0^1 dx \int_0^x dy \frac{1}{[x(d_2 - d_3) + y(d_1 - d_2) + d_3]^3}$$

$$\frac{1}{d_1 d_2 d_3 d_4}$$

$$= \int_0^1 dx \int_0^x dy \int_0^y dz \frac{\Gamma(4)}{[x(d_2 - d_1) + y(d_3 - d_2) + z(d_4 - d_3) + d_1]^4}$$

$$= \int_0^1 dx \int_0^x dy \int_0^y dz \frac{\Gamma(4)}{[x(d_3 - d_4) + y(d_2 - d_3) + z(d_1 - d_2) + d_4]^4}.$$

## 12.2 Dimensional Regularization

In high energy theory, there are 3 types of divergences: (a) ultraviolet ( $E \rightarrow \infty$ ), (b) infrared ( $E \rightarrow \epsilon \rightarrow 0+$ ), and (c) collinear divergence ( $\cos \theta \rightarrow \pm 1$ ) between a quark and a gluon:

- (a) ultraviolet divergence can be removed by renormalization with a high energy cut off or dimensional regularization;
- (b) infrared divergence can be removed with real gluon or photon emission and a low energy cut off or dimensional regularization;
- (c) collinear divergence can be removed by redefinition of parton distribution functions.

Regularization is the procedure to isolate divergences and determine the finite part for physical observables.

Example 1: For mass renormalization, we have

$$\Gamma(-1 + \epsilon) = -\left[\frac{1}{\epsilon} + 1 - \gamma + O(\epsilon)\right].$$

Example 2: For coupling renormalization, we have

$$\Gamma(\epsilon) = \frac{1}{\epsilon} - \gamma + O(\epsilon).$$

## 12.3 Mass Renormalization in $\lambda\phi^4$ Theory

In  $N$  dimensions with  $N = 4 - 2\epsilon$ , the Lagrangian density for the  $\phi^4$  theory becomes

$$\mathcal{L} = \frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{1}{2}m^2\phi^2 - \frac{\lambda}{4!}\mu^{2\epsilon}\phi^4.$$

Applying Feynman rules, we obtain the self energy as a one-loop integral

$$\begin{aligned} -i\Sigma(m^2) &= S(-i\lambda)\mu^{2\epsilon} \int \frac{d^N\ell}{(2\pi)^N} \frac{i}{\ell^2 - m^2 + i\epsilon} \\ &= \frac{1}{2}(-i\lambda)\mu^{2\epsilon} \int \frac{d^N\ell}{(2\pi)^N} \frac{i}{\ell^2 - m^2 + i\epsilon} \end{aligned}$$

where  $S$  is the symmetry factor

$$S = \frac{4 \cdot 3}{4!} = \frac{1}{2}.$$

That leads to

$$\Sigma(m^2) = \frac{1}{2}(i\lambda)\mu^{2\epsilon} \int \frac{d^N \ell}{(2\pi)^N} \frac{1}{\ell^2 - m^2 + i\epsilon}.$$

Introducing Wick rotation with  $\ell_M^0 = i\ell_E^N$  and  $\ell_M^2 = -\ell_E^2 = -\ell^2$ , we obtain

$$\begin{aligned} \Sigma(m^2) &= \frac{1}{2}(i\lambda)\mu^{2\epsilon}(2\pi)^{-N} \int id^N \ell_E \frac{-1}{\ell_E^2 + m^2 - i\epsilon} \\ &= \frac{\lambda}{2}\mu^{2\epsilon}(2\pi)^{-N} \int d^N \ell \frac{1}{\ell^2 + m^2 - i\epsilon}. \end{aligned}$$

This is similar to the following integral in the N-dimensional Euclidean space,

$$\int d^N \ell \frac{1}{(\ell^2 + M^2)^A} = \pi^{N/2} \frac{\Gamma(A - N/2)}{\Gamma(A)} \frac{1}{(M^2)^{A-N/2}}$$

with  $A = 1$  and  $M^2 = m^2 - i\epsilon$ . Thus the self energy becomes

$$\Sigma(m^2) = \frac{\lambda}{2} \mu^{2\epsilon} (2\pi)^{-N} \pi^{N/2} \frac{\Gamma(1 - N/2)}{\Gamma(1)} \frac{1}{(m^2 - i\epsilon)^{1-N/2}}.$$

Choosing  $N = 4 - 2\epsilon$ , we obtain

$$\begin{aligned} \Sigma_{\text{loop}}(m^2) &= \frac{\lambda}{32\pi^2} m^2 \Gamma(-1 + \epsilon) \left( \frac{m^2}{4\pi\mu^2} \right)^{-\epsilon} \\ &= \frac{\lambda}{32\pi^2} m^2 \left[ -\frac{1}{\epsilon} - 1 + \gamma + O(\epsilon) \right] \left( \frac{m^2}{4\pi\mu^2} \right)^{-\epsilon} \\ &= \frac{\lambda}{32\pi^2} (m^2) \left[ -\Delta_\epsilon - 1 + \log \left( \frac{m^2}{\mu^2} \right) + O(\epsilon) \right]. \end{aligned}$$

where

$$\Delta_\epsilon = \frac{1}{\epsilon} - \gamma + \ln(4\pi).$$

We have applied

$$\frac{\pi^{N/2}}{(2\pi)^N} = \frac{\pi^{N/2}}{(4\pi^2)^{N/2}} = (4\pi)^{-(N/2)} = (4\pi)^{-(2-\epsilon)} = \frac{1}{16\pi^2} (4\pi)^\epsilon,$$

as well as

$$\left(\frac{m^2}{4\pi\mu^2}\right)^{-\epsilon} = e^{\ln\left(\frac{m^2}{4\pi\mu^2}\right)^{-\epsilon}} = e^{-\epsilon \ln\left(\frac{m^2}{4\pi\mu^2}\right)} = 1 - \epsilon \ln\left(\frac{m^2}{4\pi\mu^2}\right) + \mathcal{O}(\epsilon^2),$$

and

$$\Gamma(-1 + \epsilon) = (-1) \left[ \frac{1}{\epsilon} + 1 - \gamma + \mathcal{O}(\epsilon) \right]$$

where  $\gamma$  is the Euler constant and  $\epsilon$  is an infinitesimal parameter.

## 12.4 Coupling Renormalization in $\lambda\phi^4$ Theory

In  $N$  dimensions with  $N = 4 - 2\epsilon$ , we need to define a new coupling

$$\lambda_{\text{old}} = \lambda_{\text{new}} \cdot \mu^{2\epsilon} = \lambda \mu^{2\epsilon}$$

$\lambda = \lambda_{\text{new}}$  to keep  $\lambda$  dimensionless so this theory is renormalizable.

Then the Lagrangian density becomes

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2 - \frac{\lambda}{4!} \mu^{2\epsilon} \phi^4.$$

The Next-to-Leading-Order (NLO) vertex corrections have contributions from  $s$ ,  $t$  and  $u$ -channel diagrams. Applying Feynman rules, we obtain the transition amplitude for the  $s$ -channel



contribution as

$$\begin{aligned}
 iM_1 &= \frac{1}{2}(-i\lambda)^2(\mu^{2\epsilon})^2 \int \frac{d^N \ell}{(2\pi)^N} \frac{i}{\ell^2 - m^2 + i\epsilon} \frac{i}{(\ell + q_1)^2 - m^2 + i\epsilon} \\
 &= \frac{1}{2}\lambda^2(\mu^2)^{2\epsilon}(2\pi)^{-N} \int d^N \ell \frac{1}{(\ell^2 - m^2 + i\epsilon)[(\ell + q_1)^2 - m^2 + i\epsilon]} \\
 &= \frac{1}{2}\lambda^2(\mu^2)^{2\epsilon}(2\pi)^{-N} \int d^N \ell \frac{1}{d_1 d_2}
 \end{aligned}$$

where

$$\begin{aligned}
 d_1 &= \ell^2 - m^2 + i\epsilon \\
 d_2 &= (\ell + q_1)^2 - m^2 + i\epsilon
 \end{aligned}$$

with  $q_1 = p_1 + p_2$  and  $q_1^2 = s$ .

Introducing a Feynman parameter  $x$ , we obtain

$$\begin{aligned} \frac{1}{d_1 d_2} &= \int_0^1 dx \frac{1}{[x(d_2 - d_1) + d_1]^2} \\ &= \int_0^1 dx \frac{1}{[\ell^2 + 2\ell \cdot q_1 x - m^2 + xq_1^2]^2} \end{aligned}$$

where

$$d_2 - d_1 = 2\ell \cdot q_1 + q_1^2.$$

Then the amplitude becomes

$$\begin{aligned} iM_1 &= \frac{1}{2}(\lambda)^2(\mu^2)^{2\epsilon}(2\pi)^{-N} \int d^N \ell \frac{1}{d_1 d_2} \\ &= \frac{1}{2}(\lambda)^2(\mu^2)^{2\epsilon}(2\pi)^{-N} \int d^N \ell \int_0^1 dx \frac{1}{[\ell^2 + 2\ell \cdot q_1 x - m^2 + xq_1^2]^2}. \end{aligned}$$

In  $N$  dimensions with dimensional regularization, we can switch the

order of integration and express  $M_1$  as

$$iM_1 = \frac{1}{2}(\lambda)^2(\mu^2)^{2\epsilon}(2\pi)^{-N} \int_0^1 dx \int d^N \ell \frac{1}{[\ell^2 + 2\ell \cdot q_1 x - m^2 + xq_1^2]^2}.$$

The integral in  $M_1$  can be simplified with a shift  $q = \ell + xq_1$  or  $\ell = q - xq_1$ . Then the amplitude becomes

$$iM_1 = \frac{1}{2}(\lambda)^2(\mu^2)^{2\epsilon}(2\pi)^{-N} \int_0^1 dx \int d^N q \frac{1}{[q^2 + x(1-x)q_1^2 - m^2]^2}.$$

Introducing Wick rotation in the complex  $\ell^0$  plane with  $\ell^0 = i\ell_E^N$  and  $\ell_E^N \in \mathcal{R}$ , we obtain

$$d^N \ell_M = id^N \ell_E \quad \text{and} \quad \ell_M^2 = (\ell^0)^2 - |\vec{\ell}|^2 = -(\ell^N)^2 - |\vec{\ell}|^2 = -\ell_E^2 = -\ell^2.$$

Then the s-channel amplitude becomes

$$iM_1 = i(-1)^2 \frac{1}{2}(\lambda)^2(\mu^2)^{2\epsilon}(2\pi)^{-N} \int_0^1 dx \int d^N q_E \frac{1}{[q^2 - x(1-x)q_1^2 + m^2]^2}$$

or

$$M_1 = \frac{1}{2}(\lambda)^2(\mu^2)^{2\epsilon}(2\pi)^{-N} \int_0^1 dx \int d^N q \frac{1}{[q^2 - x(1-x)q_1^2 + m^2]^2}.$$

Recall that in the  $N$ -dimensional Euclidean space, we have

$$I_N = \int d^N \ell \frac{1}{(\ell^2 + M^2)^A} = \pi^{N/2} \frac{\Gamma(A - N/2)}{\Gamma(A)} \frac{1}{(M^2)^{A - N/2}}.$$

Thus we now have

$$\int d^N q \frac{1}{[q^2 - x(1-x)q_1^2 + m^2]^2} = \int d^N q \frac{1}{[q^2 + M^2]^2} = \pi^{N/2} \frac{\Gamma(2 - N/2)}{\Gamma(2)} \frac{1}{(M^2)^{2 - N/2}}$$

with  $A = 2$  and  $M^2 = -x(1-x)q_1^2 + m^2$ .

Then the s-channel diagram becomes

$$\begin{aligned} M_1 &= \frac{1}{2}(\lambda)^2(\mu^2)^{2\epsilon}(2\pi)^{-N} \int_0^1 dx \pi^{N/2} \frac{\Gamma(2 - N/2)}{\Gamma(2)} \frac{1}{[-x(1-x)q_1^2 + m^2]^{2 - N/2}} \\ &= \frac{1}{2}(\lambda)^2(\mu^2)^{2\epsilon}(2\pi)^{-N} \pi^{N/2} \Gamma(2 - N/2) \int_0^1 dx \frac{1}{[-x(1-x)q_1^2 + m^2]^{2 - N/2}} \end{aligned}$$

with  $\Gamma(2) = 1! = 1$ .

We often choose  $N = 4 - 2\epsilon$ . Then  $2 - N/2 = \epsilon$ ,  $\Gamma(2 - N/2) = \Gamma(\epsilon)$ , and

$$\frac{\pi^{N/2}}{(2\pi)^N} = \frac{\pi^{N/2}}{(4\pi^2)^{N/2}} = (4\pi)^{-(N/2)} = (4\pi)^{-(2-\epsilon)} = \frac{1}{16\pi^2} (4\pi)^\epsilon.$$

Then the amplitude becomes

$$\begin{aligned} M_1 &= \frac{\lambda^2}{32\pi^2} \mu^{2\epsilon} (\mu^2)^\epsilon (4\pi)^\epsilon \Gamma(\epsilon) \int_0^1 dx [m^2 + x(1-x)q_1^2]^{-\epsilon} \\ &= \frac{\lambda^2}{32\pi^2} \mu^{2\epsilon} \Gamma(\epsilon) \int_0^1 dx \left[ \frac{-x(1-x)q_1^2 + m^2}{4\pi\mu^2} \right]^{-\epsilon}. \end{aligned}$$

Now let us consider

$$\Lambda^{-\epsilon} = e^{\ln(\Lambda^{-\epsilon})} = e^{-\epsilon \ln(\Lambda)} = 1 - \epsilon \ln(\Lambda) + \mathcal{O}(\epsilon^2)$$

and

$$\Gamma(-n + \epsilon) = \frac{(-1)^n}{n!} \left[ \frac{1}{\epsilon} + \psi(n+1) + \mathcal{O}(\epsilon) \right]$$

where

$$\begin{aligned}\psi(n+1) &= 1 + \frac{1}{2} + \cdots + \frac{1}{n} - \gamma, \\ \psi(1) &= -\gamma \simeq -0.5772,\end{aligned}$$

$\gamma$  is the Euler constant and  $\epsilon$  is an infinitesimal positive parameter  $\epsilon \rightarrow 0+$ . Then the amplitude becomes

$$\begin{aligned}M_1 &= \frac{\lambda^2}{32\pi^2} \mu^{2\epsilon} \left[ \frac{1}{\epsilon} - \gamma - \int_0^1 dx \ln \left( \frac{m^2 - x(1-x)q_1^2}{4\pi\mu^2} \right) + \mathcal{O}(\epsilon) \right] \\ &= \frac{\lambda^2}{32\pi^2} \mu^{2\epsilon} \left[ \frac{1}{\epsilon} - \gamma + \ln(4\pi) + \ln \left( \frac{\mu^2}{m^2} \right) - \int_0^1 dx \ln \left( 1 + \frac{-x(1-x)q_1^2}{m^2} \right) \right] \\ &= \frac{\lambda^2}{32\pi^2} \mu^{2\epsilon} \left[ \Delta_\epsilon + \ln \left( \frac{\mu^2}{m^2} \right) - \int_0^1 dx \ln \left( 1 + \frac{-q_1^2}{4m^2} 4x(1-x) \right) + \mathcal{O}(\epsilon) \right]\end{aligned}$$

where

$$\Delta_\epsilon = \frac{1}{\epsilon} - \gamma + \ln(4\pi)$$

and

$$\ln \left( \frac{m^2 - x(1-x)q_1^2}{4\pi\mu^2} \right) = -\ln(4\pi) - \ln \left( \frac{\mu^2}{m^2} \right) + \ln \left( 1 + \frac{-q_1^2}{4m^2} 4x(1-x) \right).$$

Applying the following integral

$$\int_0^1 dx \left[ 1 + \frac{4}{a}x(1-x) \right] = -2 + \sqrt{1+a} \ln \left( \frac{\sqrt{1+a}+1}{\sqrt{1+a}-1} \right) \quad \text{for } a > 0$$

we obtain

$$\int_0^1 dx \ln \left( \frac{m^2 - x(1-x)q_1^2}{4\pi\mu^2} \right) = -\ln(4\pi) - \ln \left( \frac{\mu^2}{m^2} \right) + \left[ -2 + \sqrt{1 - \frac{4m^2}{q_1^2}} \ln \frac{\sqrt{1 - \frac{4m^2}{q_1^2}} + 1}{\sqrt{1 - \frac{4m^2}{q_1^2}} - 1} \right]$$

Then the s-channel diagram with  $q_1^2 = s$  becomes

$$M_1 = \frac{\lambda^2}{32\pi^2} \mu^{2\epsilon} \left[ \frac{1}{\epsilon} - \gamma + \ln(4\pi) + 2 + \ln \left( \frac{\mu^2}{m^2} \right) - \sqrt{1 - \frac{4m^2}{s}} \ln \frac{\sqrt{1 - \frac{4m^2}{s}} + 1}{\sqrt{1 - \frac{4m^2}{s}} - 1} \right]$$

Similarly, the  $t$ - and  $u$ -channel diagrams become

$$M_2 = \frac{\lambda^2}{32\pi^2} \mu^{2\epsilon} \left[ \frac{1}{\epsilon} - \gamma + \ln(4\pi) + 2 + \ln\left(\frac{\mu^2}{m^2}\right) - \sqrt{1 - \frac{4m^2}{t}} \ln \frac{\sqrt{1 - \frac{4m^2}{t}} + \frac{4m^2}{t}}{\sqrt{1 - \frac{4m^2}{t}} - \frac{4m^2}{t}} \right]$$

and

$$M_3 = \frac{\lambda^2}{32\pi^2} \mu^{2\epsilon} \left[ \frac{1}{\epsilon} - \gamma + \ln(4\pi) + 2 + \ln\left(\frac{\mu^2}{m^2}\right) - \sqrt{1 - \frac{4m^2}{u}} \ln \frac{\sqrt{1 - \frac{4m^2}{u}} + \frac{4m^2}{u}}{\sqrt{1 - \frac{4m^2}{u}} - \frac{4m^2}{u}} \right]$$

with  $q_2^2 = (p_2 + p_3)^2 = t$  and  $q_3^2 = (p_1 + p_3)^2 = u$

Thus the total one-loop contribution at the order of  $\lambda^2$  is

$$\begin{aligned} G_4^{(1)} &= i(M_1 + M_2 + M_3) \\ &= i \frac{\lambda^2}{32\pi^2} \mu^{2\epsilon} \left[ 3 \left( \frac{1}{\epsilon} - \gamma + \ln(4\pi) \right) + F_4 \right] + \mathcal{O}(\epsilon) \\ &= i \frac{\lambda^2}{32\pi^2} \mu^{2\epsilon} [3\Delta_\epsilon + F_4] + \mathcal{O}(\epsilon). \end{aligned}$$



To the order of the  $\lambda^2$ , the 4-point vertex function or the 4-point Green function becomes

$$G_4 = G_4^{(0)} + G_4^{(1)}$$

$$= -i\lambda + i\frac{\lambda^2}{32\pi^2}\mu^{2\epsilon} [3\Delta_\epsilon + F_4] + \mathcal{O}(\epsilon) \quad \text{with}$$

$$\Delta_\epsilon = \frac{1}{\epsilon} - \gamma + \ln(4\pi), \quad \text{and}$$

$$F_4 = 6 + 3\ln\left(\frac{\mu^2}{m^2}\right) - \sqrt{1 - \frac{4m^2}{s}} \ln \frac{\sqrt{1 - \frac{4m^2}{s}} + 1}{\sqrt{1 - \frac{4m^2}{s}} - 1}$$

$$- \sqrt{1 - \frac{4m^2}{t}} \ln \frac{\sqrt{1 - \frac{4m^2}{t}} + 1}{\sqrt{1 - \frac{4m^2}{t}} - 1} - \sqrt{1 - \frac{4m^2}{u}} \ln \frac{\sqrt{1 - \frac{4m^2}{u}} + 1}{\sqrt{1 - \frac{4m^2}{u}} - 1}.$$

The divergent part

$$+i\frac{3\lambda^2}{32\pi^2}\mu^{2\epsilon} \begin{bmatrix} 1 \\ - \\ \epsilon \end{bmatrix}$$

needs to be removed by renormalization with counter terms.

## Loop Integrals in the Minkowski Space

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{1}{(\ell^2 + 2\ell \cdot p + M^2)^A} = \frac{\Gamma(A - N/2)}{\Gamma(A)} \frac{1}{(M^2 - p^2)^{A - N/2}} \quad (9)$$

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{l_\mu}{(\ell^2 + 2\ell \cdot p + M^2)^A} = \frac{\Gamma(A - N/2)}{\Gamma(A)} \frac{-p_\mu}{(M^2 - p^2)^{A - N/2}} \quad (10)$$

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{l_\mu l_\nu}{(\ell^2 + 2\ell \cdot p + M^2)^A} = \frac{1}{\Gamma(A)}$$

$$\times \left[ \frac{1}{2} \eta_{\mu\nu} \frac{\Gamma(A-1-N/2)}{(M^2-p^2)^{A-1-N/2}} + p_\mu p_\nu \frac{\Gamma(A-N/2)}{(M^2-p^2)^{A-N/2}} \right]$$

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{l_\mu l_\nu l_\rho}{(\ell^2 + 2\ell \cdot p + M^2)^A} = \frac{-1}{\Gamma(A)}$$

$$\times \left[ \frac{1}{2} (\eta_{\mu\nu} p_\rho + \eta_{\mu\rho} p_\nu + \eta_{\nu\rho} p_\mu) \frac{\Gamma(A-1-N/2)}{(M^2-p^2)^{A-1-N/2}} + p_\mu p_\nu p_\rho \frac{\Gamma(A-N/2)}{(M^2-p^2)^{A-N/2}} \right]$$

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{l_\mu l_\nu l_\rho l_\sigma}{(\ell^2 + 2\ell \cdot p + M^2)^A} = \frac{1}{\Gamma(A)}$$

$$\times \left\{ \frac{1}{4} (\eta_{\mu\nu} \eta_{\rho\sigma} + \eta_{\mu\rho} \eta_{\nu\sigma} + \eta_{\nu\rho} \eta_{\mu\sigma}) \frac{\Gamma(A-2-N/2)}{(M^2-p^2)^{A-2-N/2}} \right.$$

$$+ \frac{1}{2} (\eta_{\mu\nu} p_\rho p_\sigma + \eta_{\mu\rho} p_\nu p_\sigma + \eta_{\mu\sigma} p_\nu p_\rho + \eta_{\nu\rho} p_\mu p_\sigma + \eta_{\nu\sigma} p_\mu p_\rho + \eta_{\rho\sigma} p_\mu p_\nu) \frac{\Gamma(A-N/2)}{(M^2-p^2)^{A-N/2}}$$

$$\left. + p_\mu p_\nu p_\rho p_\sigma \frac{\Gamma(A-N/2)}{(M^2-p^2)^{A-N/2}} \right\}$$

All these integrals can be evaluated with dimensional regularization,  
 $N = 4 - 2\epsilon$  and  $\epsilon$  is infinitesimal.

## 10.1 12.5 Large Mass Expansion

For  $x^2 < 1$ , we have

$$(1 \pm x)^{-1} = 1 \mp x + x^2 + O(x^3). \quad (14)$$

For  $(q^2/M^2)^2 < 1$ , we can expand the propagator<sup>a</sup> as

$$\begin{aligned} [(l+q)^2 + M^2]^{-1} &= [\ell^2 + M^2 + 2\ell \cdot q + q^2]^{-1} \\ &= [\ell^2 + M^2]^{-1} \left[ 1 + \frac{2\ell \cdot q + q^2}{\ell^2 + M^2} \right]^{-1} \\ &= [\ell^2 + M^2]^{-1} \left[ 1 - \frac{2\ell \cdot q + q^2}{\ell^2 + M^2} + \frac{(2\ell \cdot q + q^2)^2}{(\ell^2 + M^2)^2} + O(M^{-6}) \right] \end{aligned} \quad (15)$$

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<sup>a</sup>M. Veltman, Acta Physica Polonica **B8** (1977) 493.

## Three-point Scalar Integral

Now let us calculate the 3-point scalar integral

$$\begin{aligned}C_0 &= \int \frac{d^N}{(2\pi)^N} \frac{1}{d_1 d_2 d_3} \\d_1 &= \ell^2 - m^2 + i\epsilon \\d_2 &= (\ell + p_1)^2 + i\epsilon \\d_3 &= (\ell + p_1 + p_2)^2 - m^2 + i\epsilon\end{aligned}$$

with  $p_1^2 = m^2 = p_2^2$  and  $k^2 = (p_1 + p_2)^2 = s$  as an integral in  $x$  and  $y$ .

Let us introduce Feynman parameters for the denominator of  $C_0$ :

$$\frac{1}{d_1 d_2 d_3} = \Gamma(3) \int_0^1 dx \int_0^x dy \frac{1}{[x(d_2 - d_1) + y(d_3 - d_2) + d_1]^3}$$

where

$$d_2 - d_1 = 2\ell \cdot p_1 + p_1^2 + m^2,$$

$$d_3 - d_2 = 2(\ell + p_1) \cdot p_2 + p_2^2 - m^2 = 2\ell \cdot p_2 + 2p_1 \cdot p_2 + p_2^2 - m^2.$$

Then the denominator becomes

$$\begin{aligned} x(d_2 - d_1) + y(d_3 - d_2) + d_1 &= \ell^2 + 2\ell \cdot (xp_1 + yp_2) \\ &\quad + x(p_1^2 + m^2) + y(2p_1 \cdot p_2 + p_2^2 - m^2) - m^2 + \end{aligned}$$

Applying dimensional regularization, we can switch the order of the  $\ell$  and the  $x, y$  integrals and express  $C_0$  as

$$\begin{aligned} C_0 &= \frac{\Gamma(3)}{(2\pi)^N} \int_0^1 dx \int_0^x dy \times \\ &\quad \int \frac{d^N \ell}{[\ell^2 + 2\ell \cdot (xp_1 + yp_2) + x(p_1^2 + m^2) + y(2p_1 \cdot p_2 + p_2^2 - m^2) - m^2 + i\epsilon]} \end{aligned}$$

Let us shift the integration momentum with  $q = \ell + xp_1 + yp_2$  or



$\ell = q - xp_1 - yp_2$ . Then  $C_0$  becomes

$$C_0 = \frac{\Gamma(3)}{(2\pi)^N} \int_0^1 dx \int_0^x dy \int d^N q \frac{1}{[q^2 - M^2]^3}$$

where

$$\begin{aligned} M^2 &= -[-(xp_1 + yp_2)^2 + x(p_1^2 + m^2) + y(2p_1 \cdot p_2 + p_2^2 - m^2) - m^2 + i\epsilon] \\ &= p_1^2 x^2 + p_2^2 y^2 + 2p_1 \cdot p_2 xy - x(p_1^2 + m^2) - y(2p_1 \cdot p_2 + p_2^2 - m^2) + m^2 - \end{aligned}$$

Introducing Wick rotation with  $q^0 = iq_E^N$  and  $q_M^2 = -q_E^2 = -q^2$ , we obtain

$$C_0 = i(-1)^3 \frac{\Gamma(3)}{(2\pi)^N} \int_0^1 dx \int_0^x dy \int d^N q \frac{1}{[q^2 + M^2]^3}$$

in the Euclidean space. Recall that in  $N$  dimensional Euclidean space

$$\int d^N \ell \frac{1}{(\ell^2 + M^2)^A} = \pi^{N/2} \frac{\Gamma(A - N/2)}{\Gamma(A)} \frac{1}{(M^2)^{A - N/2}}.$$

For  $A = 3$ , the 3-point scalar integral becomes

$$\begin{aligned}
 C_0 &= -i \frac{\Gamma(3)}{(2\pi)^N} \int_0^1 dx \int_0^x dy \pi^{N/2} \frac{\Gamma(3 - N/2)}{\Gamma(3)} \frac{1}{(M^2)^{3-N/2}} \\
 &= -i \frac{\pi^{N/2}}{(2\pi)^N} \Gamma(3 - N/2) \int_0^1 dx \int_0^x dy \frac{1}{(M^2)^{3-N/2}}
 \end{aligned}$$

where

$$M^2 = p_1^2 x^2 + p_2^2 y^2 + 2p_1 \cdot p_2 xy - x(p_1^2 + m^2) - y(2p_1 \cdot p_2 + p_2^2 - m^2) + m^2 - i\epsilon.$$

Choosing  $N = 4 - 2\epsilon$ , we obtain

$$\Gamma\left(3 - \frac{N}{2}\right) = \Gamma(1 + \epsilon) = 1 + \mathcal{O}(\epsilon).$$

Since it is finite, we may set  $N = 4$ . Then  $C_0$  becomes

$$C_0 = \frac{-i}{16\pi^2} \int_0^1 dx \int_0^x dy \frac{1}{M^2}$$

where

$$M^2 = p_1^2 x^2 + p_2^2 y^2 + 2p_1 \cdot p_2 xy - x(p_1^2 + m^2) - y(2p_1 \cdot p_2 + p_2^2 - m^2) + m^2 - i\epsilon.$$

For  $p_1^2 = m^2 = p^2$  and  $s = (p_1 + p_2)^2 = 2p_1 \cdot p_2 + 2m^2$ , we have

$$M^2 = m^2 x^2 + m^2 y^2 + (s - 2m^2)xy - 2m^2 x - (s - 2m^2)y + m^2 - i\epsilon.$$

In general,  $C_0$  is expressed in terms of Spence functions.

## 10.2 The Gamma Functions

The following formulae are very useful for dimensional regularization.

$$\begin{aligned}\Gamma(-n + \epsilon) &= \frac{(-1)^n}{n!} \left[ \frac{1}{\epsilon} + \psi_1(n + 1) + O(\epsilon) \right] \\ \Gamma(\epsilon) &= \frac{1}{\epsilon} - \gamma + O(\epsilon) \\ \Gamma(-1 + \epsilon) &= -\left[ \frac{1}{\epsilon} + 1 - \gamma + O(\epsilon) \right] \\ \Gamma(-2 + \epsilon) &= \frac{1}{2} \left[ \frac{1}{\epsilon} + 1 + \frac{1}{2} - \gamma + O(\epsilon) \right]\end{aligned}\tag{16}$$

where  $\gamma$  is the Euler constant, and

$$\begin{aligned}\psi_1(z) &= \frac{d[\ln \Gamma(z)]}{dz} = \frac{\Gamma'(z)}{\Gamma(z)} \\ \psi_1(n + 1) &= 1 + \frac{1}{2} + \dots + \frac{1}{n} - \gamma\end{aligned}\tag{17}$$

# Divergence Cancellation

Here are some identities for divergence cancellation:

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{l_\mu l_\nu}{(\ell^2 + M^2)^A} = \frac{1}{\Gamma(A)} \frac{1}{2} \eta_{\mu\nu} \frac{\Gamma(A - 1 - N/2)}{(M^2)^{A-1-N/2}}$$

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{l_\mu l_\nu l_\rho l_\sigma}{(\ell^2 + M^2)^A} = \frac{1}{\Gamma(A)} \frac{1}{4} (\eta_{\mu\nu} \eta_{\rho\sigma} + \eta_{\mu\rho} \eta_{\nu\sigma} + \eta_{\nu\rho} \eta_{\mu\sigma}) \frac{\Gamma(A - 2 - N/2)}{(M^2)^{A-2-N/2}}$$

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{1}{(\ell^2 + M^2)^2} = \frac{1}{\Gamma(2)} \frac{\Gamma(2 - N/2)}{(M^2)^{2-N/2}}$$

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{l_\mu l_\nu}{(\ell^2 + M^2)^3} = \frac{1}{4} \eta_{\mu\nu} \int \frac{d^N \ell}{i\pi^{N/2}} \frac{(\ell^2 + M^2)}{(\ell^2 + M^2)^3}$$

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{l_\mu l_\nu \ell^2}{(\ell^2 + M^2)^4} = \frac{1}{4} \eta_{\mu\nu} \int \frac{d^N \ell}{i\pi^{N/2}} \frac{(\ell^2 + M^2)^2}{(\ell^2 + M^2)^4} - M^2 \int \frac{d^N \ell}{i\pi^{N/2}} \frac{l_\mu l_\nu}{(\ell^2 + M^2)^4}$$

$$\int \frac{d^N \ell}{i\pi^{N/2}} \frac{l_\mu l_\nu l_\rho l_\sigma}{(\ell^2 + M^2)^4} = \frac{1}{\Gamma(4)} \frac{1}{4} (\eta_{\mu\nu} \eta_{\rho\sigma} + \eta_{\mu\rho} \eta_{\nu\sigma} + \eta_{\nu\rho} \eta_{\mu\sigma}) \int \frac{d^N \ell}{i\pi^{N/2}} \frac{(\ell^2 + M^2)}{(\ell^2 + M^2)^4}$$

## 10.3 Dirac Matrices

Useful Formulae:

$$\gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu = 2\eta^{\mu\nu}$$

$$\eta_{\mu\nu} \eta^{\mu\nu} = N$$

$$\gamma_\mu \not{a} \gamma^\mu = -2(1 - \epsilon) \not{a}$$

$$\gamma_\mu \not{a} \not{b} \gamma^\mu = 4a \cdot b - 2\epsilon \not{a} \not{b}$$

$$\gamma_\mu \not{a} \not{b} \not{c} \gamma^\mu = -2 \not{c} \not{b} \not{a} + 2\epsilon \not{a} \not{b} \not{c}$$

where  $N = 4 - 2\epsilon$ .