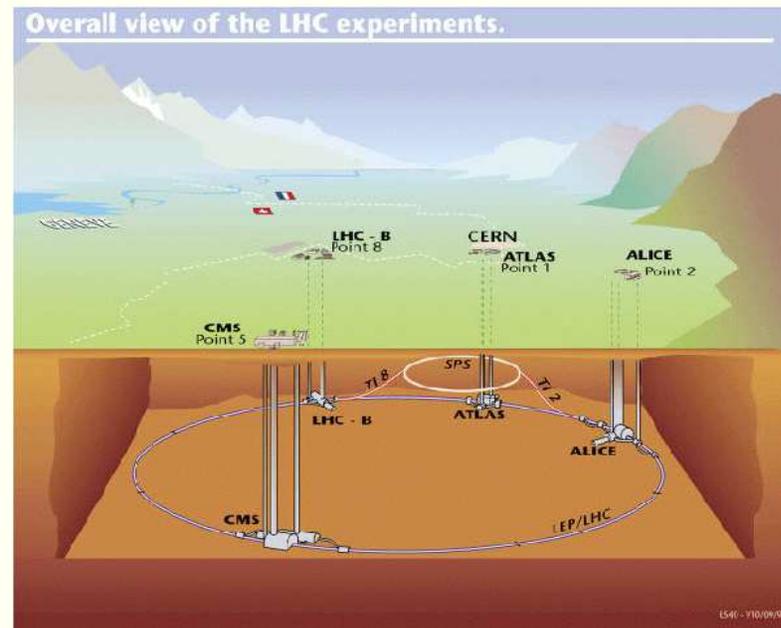


Beyond the Standard Model

Howard Baer

University of Oklahoma

- ★ neutrinos
- ★ axions
- ★ GUTs
- ★ extra dimensions: LED, RS
- ★ little Higgs
- ★ supersymmetry



The Standard Model of Particle Physics

- ★ gauge symmetry: $SU(3)_C \times SU(2)_L \times U(1)_Y \Rightarrow g, W^\pm, Z^0, \gamma$
- ★ matter content: 3 generations quarks and leptons

$$\begin{pmatrix} u \\ d \end{pmatrix}_L, u_R, d_R; \begin{pmatrix} \nu \\ e \end{pmatrix}_L, e_R \quad (1)$$

- ★ Higgs sector \Rightarrow spontaneous electroweak symmetry breaking:

$$\phi = \begin{pmatrix} \phi^+ \\ \phi_0 \end{pmatrix} \quad (2)$$

- ★ \Rightarrow massive W^\pm, Z^0 , quarks and leptons
- ★ $\mathcal{L} = \mathcal{L}_{gauge} + \mathcal{L}_{matter} + \mathcal{L}_{Yuk.} + \mathcal{L}_{Higgs}$: 19 parameters
- ★ good-to-excellent description of (almost) *all* accelerator data!

Shortcomings of SM

Data

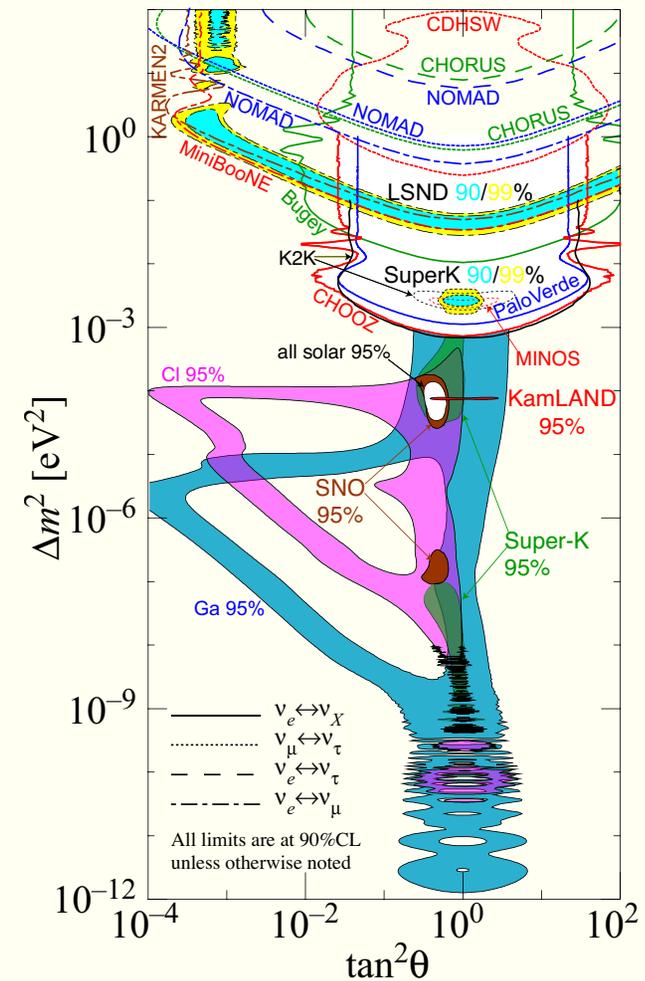
- ★ neutrino masses and mixing
- ★ baryogenesis (matter anti-matter asymmetry)
- ★ cold dark matter
- ★ dark energy

Theory

- ★ quadratic divergences in scalar sector \Rightarrow fine-tuning
- ★ origin of generations
- ★ explanation of masses/ mixing angles
- ★ origin of gauge symmetry/ quantum numbers
- ★ unification with gravity

The neutrino revolution: 1999 \Rightarrow today

- ★ Solar ν problem: Davis-Bahcall
- ★ confirmed: Gallex, SAGE
- ★ SuperK atmos. oscillations
- ★ SNO ν NC observations
- ★ MINOS
- ★ KamLand
- ★ *K2K*



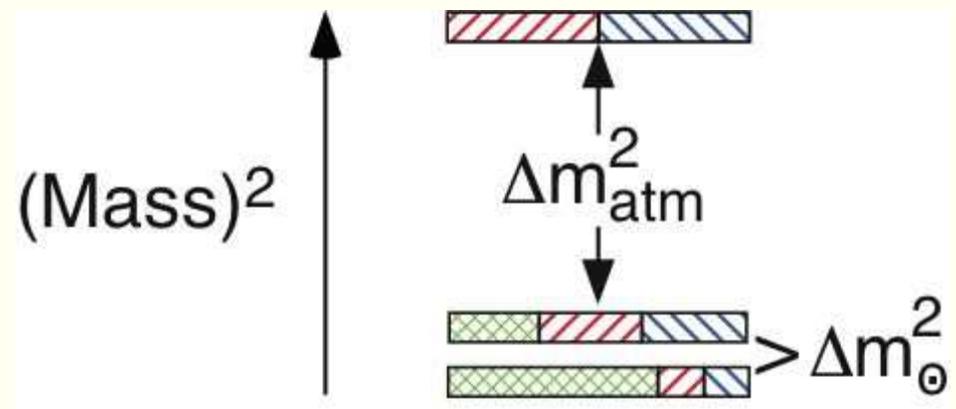
Interpretation: massive ν s and oscillations

★ SuperK/*K2K*/Minos atm osc.:

- ν s oscillate: are massive!
- $\Delta m_{atm.}^2 \sim 0.003 \text{ eV}^2$
- $\theta_{atm} \sim 45^\circ$

★ Solar ν problem

- MSW matter oscillations
- $\nu_e(r=0) \Rightarrow \nu_2(r=R_\odot)$
- measure ν_e content of ν_2
- $\Delta m_\odot^2 \sim 5 \times 10^{-5} \text{ eV}^2$
- $\theta_\odot \sim 30^\circ$

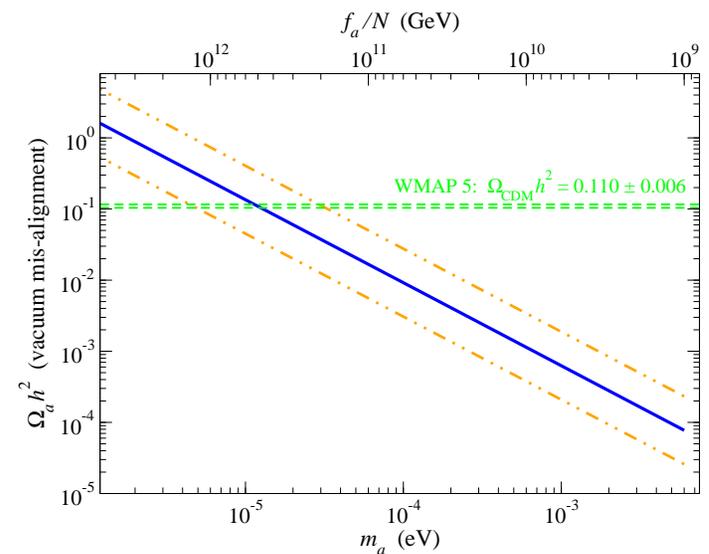


ν questions

- ★ what is absolute m_{ν_i} ?
- ★ is ν spectra normal or inverted?
- ★ are ν s Majorana or Dirac?
- ★ what is θ_{13} ?
- ★ CP violation in PMNS mixing matrix?
- ★ interpretation in terms of see-saw:
 - $m_{\nu_i} \sim (f_{\nu_i} v)^2 / M_{N_i}$
- ★ what is M_{N_i} scale?
- ★ implications for GUTs
 - disparate scales: m_{weak} and $M_{N_i} \sim M_{GUT}$?
- ★ implications for leptogenesis

Axions

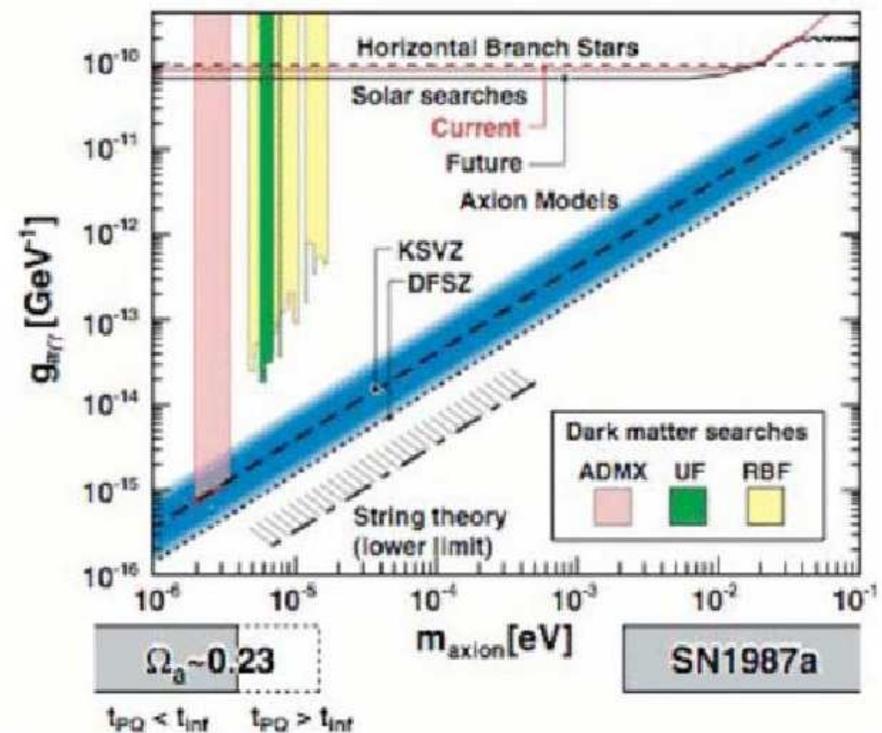
- ★ $\mathcal{L}_{QCD} \ni \frac{g_s^2 \bar{\theta}}{32\pi^2} F_A^{\mu\nu} \tilde{F}_{A\mu\nu}$: expect $\bar{\theta} \sim 1$ but $\bar{\theta} \lesssim 10^{-9}$
- ★ Peccei-Quinn solution to strong CP problem in QCD: $a(x)$
- ★ pseudo-Goldstone boson from
PQ breaking at scale $f_a \sim 10^9 - 10^{12}$ GeV
- ★ non-thermally produced
via vacuum mis-alignment as *cold* DM
 - $m_a \sim \Lambda_{QCD}^2 / f_a \sim 10^{-6} - 10^{-1} eV$
 - $\Omega_a h^2 \sim \frac{1}{2} \left[\frac{6 \times 10^{-6} eV}{m_a} \right]^{7/6} h^2$
 - astro bound: stellar cooling $\Rightarrow m_a < 10^{-5}$
 - a couples to EM field: $a - \gamma - \gamma$ coupling (Sikivie)



Axion microwave cavity searches

★ ongoing searches: ADMX experiment

- Livermore \Rightarrow U Wash.
- Phase I: probe KSVZ
for $m_a \sim 10^{-6} - 10^{-5} \text{ eV}$
- Phase II: probe DFSZ
for $m_a \sim 10^{-6} - 10^{-5} \text{ eV}$
- beyond Phase II:
probe higher values m_a



Grand unified theories: $SU(5)$

- ★ unify $SU(3)_C \times SU(2)_L \times U(1)_Y$ into $SU(5)$: single coupling constant: g_5
 - $SU(5) \rightarrow SU(3)_C \times SU(2)_L \times U(1)_Y$ at $Q = M_{GUT} \sim 10^{15}$ GeV
 - new X_μ, Y_μ gauge bosons
 - predict p -decay: $\tau_p \sim 10^{30}$ years
 - experiment: $\tau_p \gtrsim 10^{33}$ years (rule out simplest models)
- ★ partial matter unification
 - $d_R^c, \epsilon^{ab} L_b \ni \mathbf{5}^*$ of $SU(5)$
 - $Q_a, u_R^c, e_R^c \ni \mathbf{10}$ of $SU(5)$
 - * Charge quantization: $Q(e^-) = 3Q(d_R)$!
- ★ $m_b \sim 3m_\tau$ (true)
- ★ doublet-triplet splitting: $\mathbf{5}_H \ni \phi_{SM}, H_c$
- ★ gauge hierarchy problem: $m_H \rightarrow M_{GUT}$ fine-tune $1 : 10^{28}$

Grand unified theories: $SO(10)$

- ★ unify $SU(3)_C \times SU(2)_L \times U(1)_Y$ into $SO(10)$
 - 45 gauge bosons
- ★ complete matter unification of each generation
 - conjugate spinor reps **16** and **16***
 - **16** \ni **10** + **5*** + **1** of $SU(5)$
 - * all matter unified!
 - * **1** occupied by ν_R
 - * $SO(10)$ breaking $\Rightarrow M_{N_i}$ (Majorana): see-saw m_ν
- ★ explain anomaly cancellation for SM and $SU(5)$
- ★ still suffers from gauge hierarchy problem
- ★ main idea probably right; implementation probably wrong

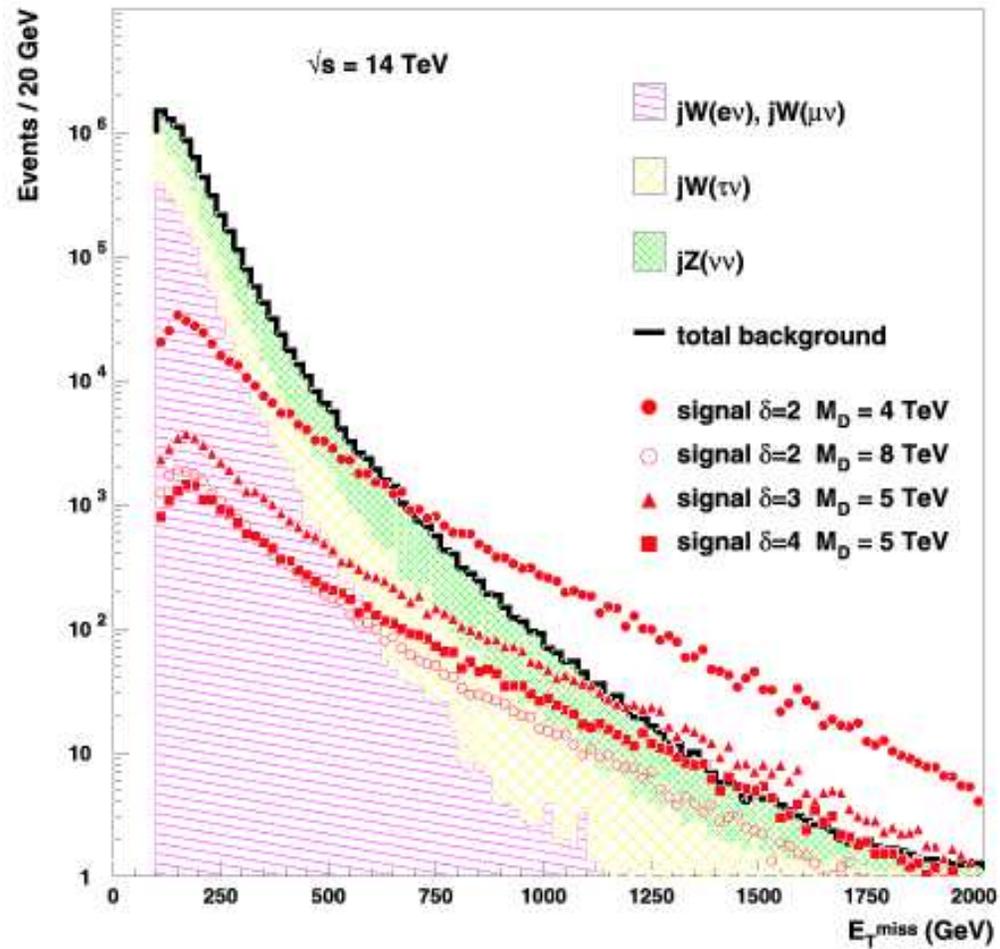
Large extra dimensions (LED or ADD)

- ★ assume SM lies on 3-brane surface embedded in $d = 4 + n$ -dim'l bulk
- ★ fundamental scale in d -dim's: $M_* = ?$
- ★ fundamental scale in 3-brane: $M_P = 2.4 \times 10^{18}$ GeV
- ★ compactify n -dim's on *e.g.* a torus
- ★ $M_P^2 = M_*^{n+2} V_n$ where V_n is volume of extra dim's
- ★ for $n > 2$, then $M_* \sim m_{weak}$ allowed! Sol'n to hierarchy problem? But then why $r_{Pl} \ll r_c$?

LED phenomenology

- ★ compactification \Rightarrow KK-excited gravitons $g_{\mu\nu}^n$: only $g_{\mu\nu}^0$ is massless
- ★ couple to matter $\sim 1/M_P$
- ★ KK-graviton mass splitting: tiny
- ★ tiny graviton emission cross-section compensated by vast number of graviton emissions possible: M_P factors cancel!
- ★ can see *e.g.* $pp \rightarrow g + g_{\mu\nu}^n$ at LHC!
- ★ can observe quantum gravity at colliders!
- ★ also, black hole production?

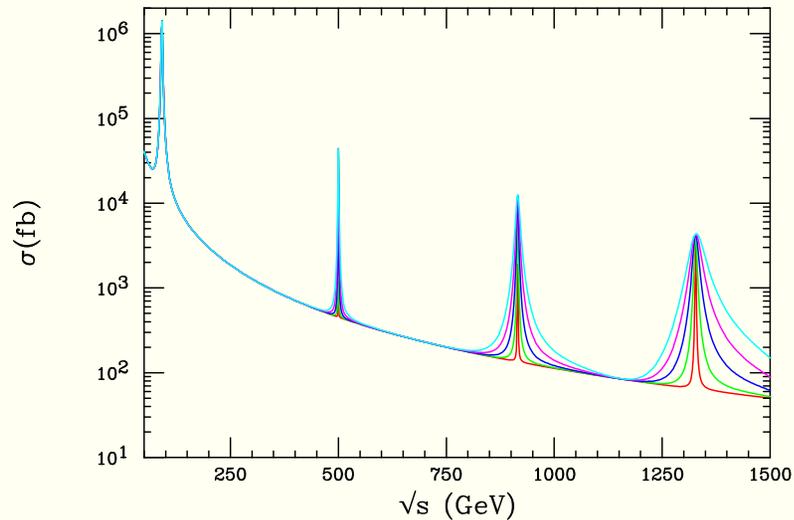
KK-graviton prod'n at LHC



Warped extra dimensions (Randall-Sundrum RS)

- ★ assume SM and hidden branes separated in warped (AdS_5) spacetime
- ★ warped metric: $ds^2 = e^{-2kr_c\phi}\eta_{\mu\nu}dx^\mu dx^\nu + r_c^2 d\phi^2$
- ★ solve Einstein eq'ns in 5-d; compactify 5th dim'n on an orbifold
- ★ find $m_{weak} = e^{-kr_c\pi} M_P$
- ★ weak scale emerges as exponential suppression of Planck scale
- ★ consequence: KK-excited gravitons with weak scale coupling and TeV-scale mass splitting!
- ★ can see excited graviton resonances at colliders!

RS phenomenology



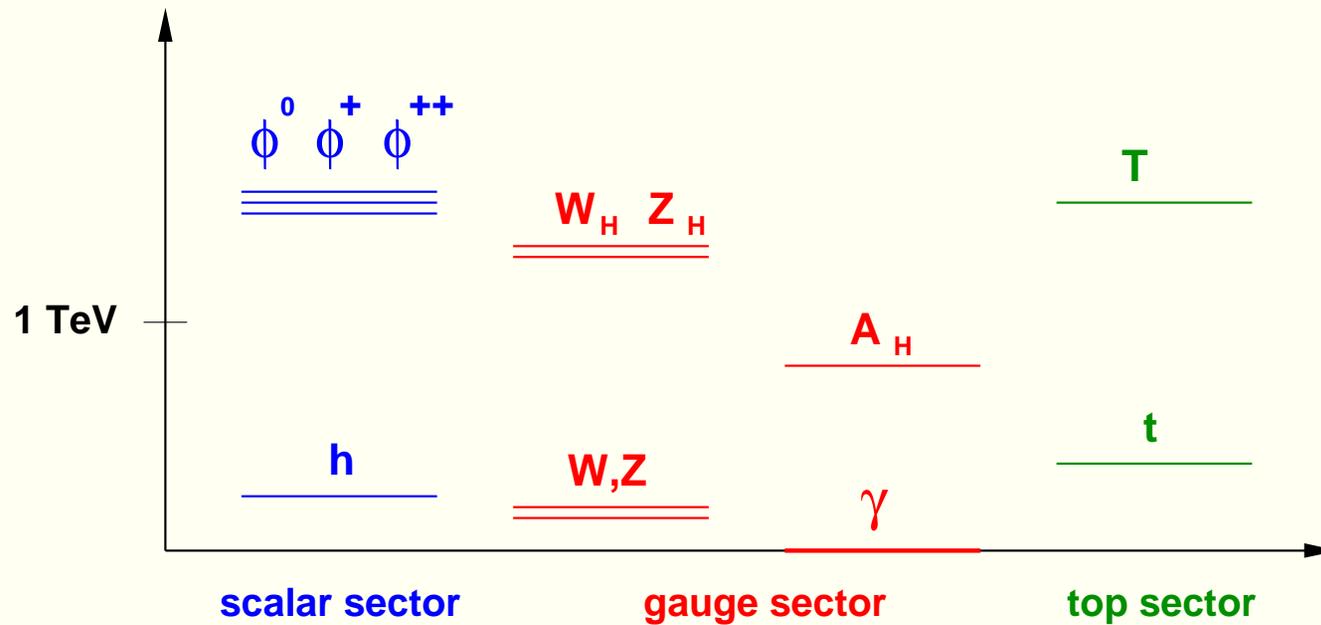
Davoudiasl, Hewett, Rizzo

- KK -excited gravitons: weak scale mass spacing via zeros of Bessel functions
- can produce RS gravitons at e^+e^- or hadron colliders
- massive graviton resonances couple to all matter pairs

Little Higgs theories

- New approach to EWSB: Arkani-Hamed, Cohen, Georgi, 2001
- Higgs field arises as pseudo-Nambu-Goldstone boson from “collective” symmetry breaking
- Symmetry \Rightarrow quadratic divergences to m_H^2 cancel at 1-loop (2-loop and higher quad. divergences remain)
- Natural cut-off of theory is ~ 10 TeV to avoid “little hierarchy problem”
- All LH theories predict new particles at 1-10 TeV scale
 - new gauge bosons A_H, W_H^\pm, W_H^0 to cancel gauge boson loops in m_H^2
 - new top partner fermions T to cancel top loop in m_H^2
 - new scalars to cancel Higgs self coupling loops
- precise details model-dependent: most popular: littlest Higgs with $SU(5)/SO(5)$

Spectrum from Little Higgs theories



T-parity in Little Higgs theories (LHT)

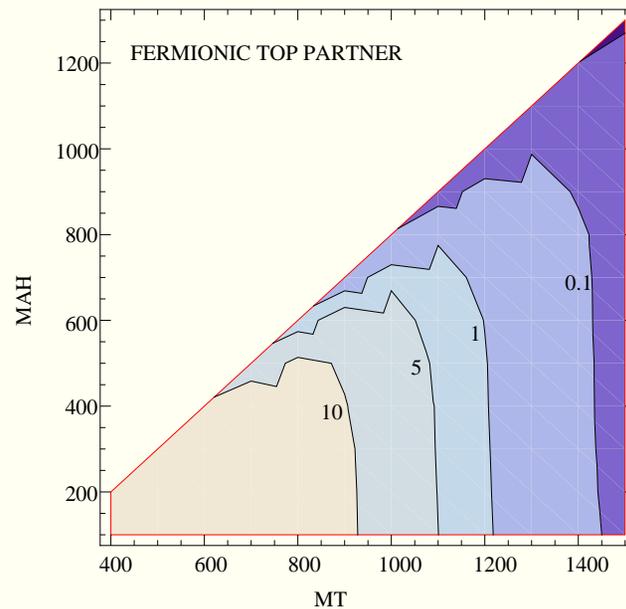
- It was found that LH models tend to give large corrections to precision EW observables unless $m_{LH} \rightarrow 10$ TeV
- This re-introduces fine-tunings in Higgs sector
- EWPOs can be saved by introducing T -parity (Cheng and Low)
 - SM particles: t -even
 - new GBs, scalars, some top-partners: t -odd
 - then contributions to EWPOs only occur at loop level
 - can allow much lighter new particle states
- t -odd particles produced in pairs
- t odd particles decay to other t -odd states
- Lightest t -odd particle absolutely stable: DM candidate, usually A_H (but see Hill+Hill anomalies paper)

LHT models at LHC

- main search channel: $pp \rightarrow T\bar{T}$ with $T \rightarrow tA_H$
- search for $t\bar{t} + \cancel{E}_T$ states
- see Cheng, Low and Wang, PRD74, 055001 (2006)
- Matsumoto, Nojiri, Nomura, PRD75, 055006 (2007)
- Belyaev, Chen, Tobe and Yuan, PRD74, 115020 (2006)
- Carena, Hubisz, Perelstein, Verdier, PRD75, 091701 (2007)
- Han, Mahbubani, Walker, Wang, arXiv:0803.3820 (2008)

$T\bar{T}$ discovery at LHC

- Han, Mahbubani, Walker, Wang, arXiv:0803.3820 (2008)
- significance after cuts with 100 fb^{-1} at LHC



Supersymmetry (SUSY)

- ★ This symmetry is similar to non-Abelian gauge symmetry except that:
 - transformation is $e^{i\bar{\alpha}Q}$, where Q is a (Majorana) spinor generator, and α is a spinorial set of parameters with $\bar{\alpha} = \alpha^\dagger \gamma_0$
 - SUSY transforms bosons \Leftrightarrow fermions
 - SUSY is a *spacetime* symmetry: the “square-root” of a translation
 - action is invariant under SUSY, but not Lagrangian (total derivative)
- ★ Can construct SUSY gauge theories
- ★ Can construct (softly broken) SUSY SM: MSSM
- ★ Solves problem of SM scalar fields: cancellation of quadratic divergences
- ★ allows for stable theories with vastly different mass scales: *e.g.* $M_{weak} \sim 10^3$ GeV and $M_{GUT} \sim 10^{16}$ GeV
- ★ *local* SUSY where $\alpha(x)$ spacetime dependent: supergravity and GR (but non-renormalizable; go to string theory?)

Minimal Supersymmetric Standard Model (MSSM)

- ★ Adopt gauge symmetry of Standard Model: $SU(3)_C \times SU(2)_L \times U(1)_Y$
 - gauge boson plus spin $\frac{1}{2}$ gaugino \in gauge superfield
- ★ SM fermions \in chiral scalar superfields: \Rightarrow scalar partner for each SM fermion helicity state
 - electron $\Leftrightarrow \tilde{e}_L$ and \tilde{e}_R
- ★ *two* Higgs doublets to cancel triangle anomalies: H_u and H_d
- ★ add all admissible soft SUSY breaking terms
- ★ resultant Lagrangian has 124 parameters!
- ★ Lagrangian yields mass eigenstates, mixings, Feynman rules for scattering and decay processes
- ★ predictive model!

Physical states of MSSM:

- ★ usual SM gauge bosons, quarks and leptons
- ★ gluino: \tilde{g}
- ★ bino, wino, neutral higgsinos \Rightarrow neutralinos: $\tilde{Z}_1, \tilde{Z}_2, \tilde{Z}_3, \tilde{Z}_4$
- ★ charged wino, higgsino \Rightarrow charginos: $\tilde{W}_1^\pm, \tilde{W}_2^\pm$
- ★ squarks: $\tilde{u}_L, \tilde{u}_R, \tilde{d}_L, \tilde{d}_R, \dots, \tilde{t}_1, \tilde{t}_2$
- ★ sleptons: $\tilde{e}_L, \tilde{e}_R, \tilde{\nu}_e, \dots, \tilde{\tau}_1, \tilde{\tau}_2, \tilde{\nu}_\tau$
- ★ Higgs sector enlarged: h, H, A, H^\pm
- ★ a plethora of new states to be found at LHC!

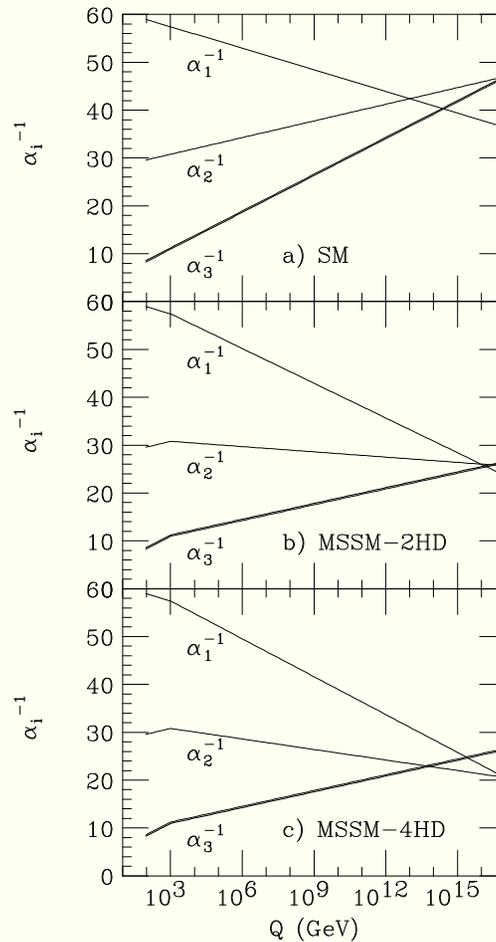
Supergravity (SUGRA)

- ★ $e^{i\bar{\alpha}Q}$ with $\alpha(x)$: *local* SUSY transformation
 - forces introduction of spin 2 graviton and spin $\frac{3}{2}$ gravitino
 - resultant theory \Rightarrow General Relativity in classical limit!
- ★ rules for Lagrangian in supergravity gauge theory: Cremmer et al. (1983)
- ★ fertile ground: supergravity \cup grand unification: LE limit of superstring?
- ★ minimal supergravity model (mSUGRA)
- ★ $m_0, m_{1/2}, A_0, \tan\beta, \text{sign}(\mu)$
 - m_0 = mass of all scalars at $Q = M_{GUT}$
 - $m_{1/2}$ = mass of all gauginos at $Q = M_{GUT}$
 - A_0 = trilinear soft breaking parameter at $Q = M_{GUT}$
 - $\tan\beta$ = ratio of Higgs vevs
 - μ = SUSY Higgs mass term; magnitude determined by REWSB!

Some successes of SUSY GUT theories

- ★ SUSY divergence cancellation maintains hierarchy between GUT scale $Q = 10^{16}$ GeV and weak scale $Q = 100$ GeV
- ★ gauge coupling unification!

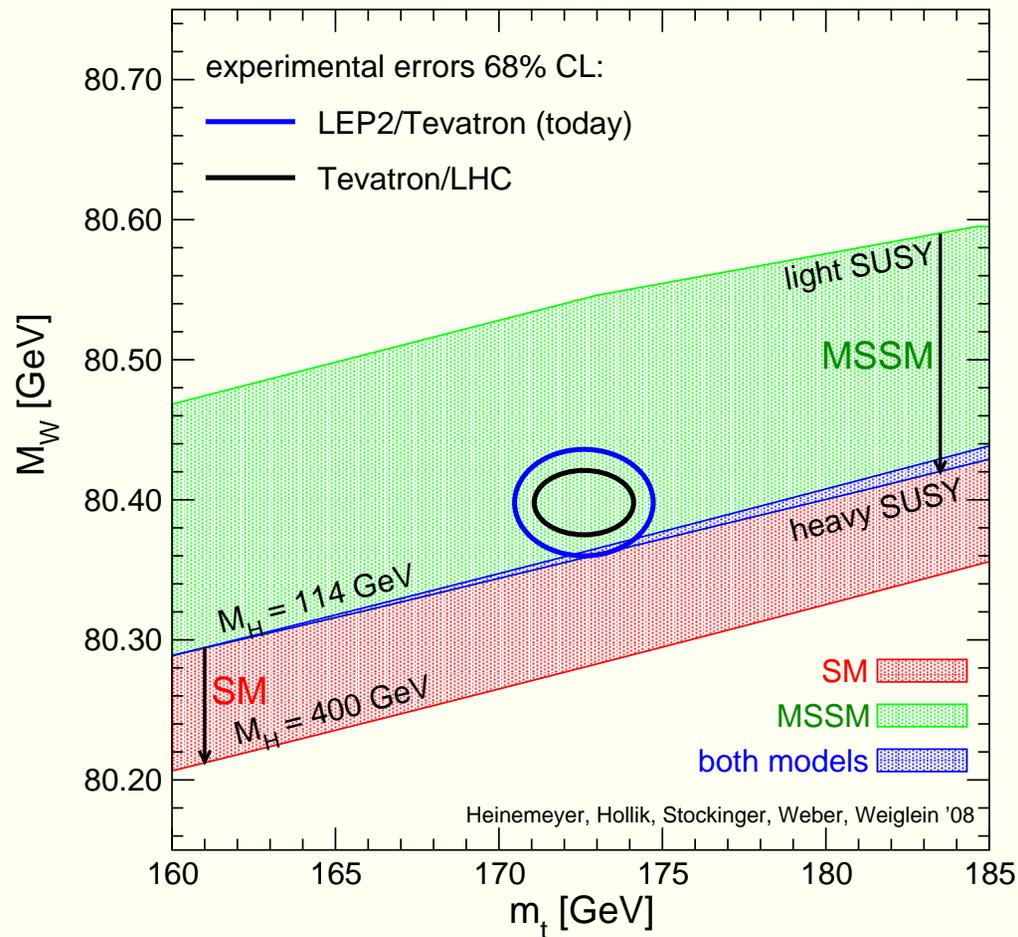
Gauge coupling evolution



Some successes of SUSY GUT theories

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- ★ gauge coupling unification!
- ★ Lightest Higgs mass $m_h \lesssim 135$ GeV as indicated by radiative corrections!

Precision electroweak data and the Higgs mass:

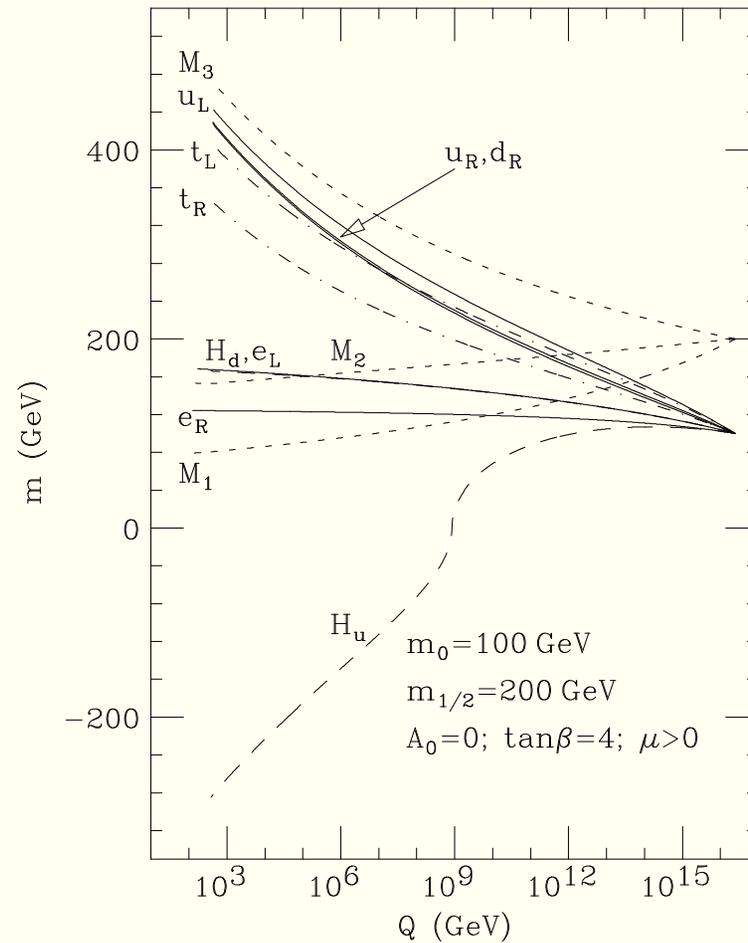


S. Heinemeyer et al.

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- ★ gauge coupling unification!
- ★ Lightest Higgs mass $m_h \lesssim 130$ GeV as indicated by radiative corrections!
- ★ radiative breaking of EW symmetry if $m_t \sim 100 - 200$ GeV!

Soft term evolution and radiative EWSB

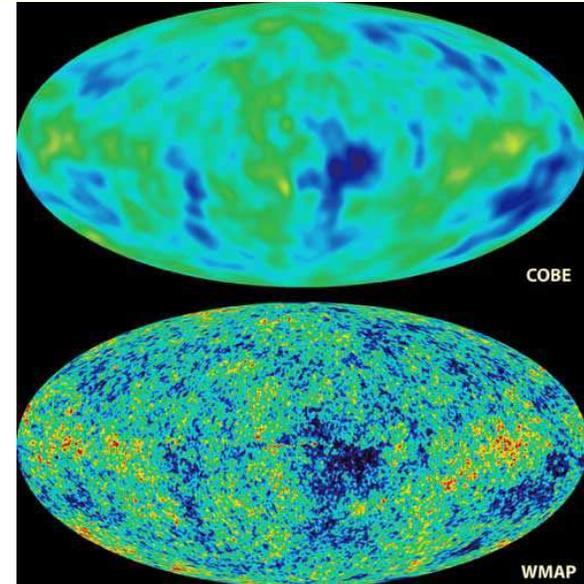


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- ★ Lightest Higgs mass $m_h \lesssim 130$ GeV as indicated by radiative corrections!
- ★ radiative breaking of EW symmetry if $m_t \sim 100 - 200$ GeV!
- ★ dark matter candidate: lightest neutralino \tilde{Z}_1
- ★ stabilize neutrino see-saw scale vs. weak scale
- ★ $SO(10)$ SUSY GUT: baryogenesis via leptogenesis
- ★ can give dark energy via CC Λ (but need huge fine-tuning...)
 - SUGRA = low energy limit of superstring?
 - stringy multiverse: anthropic selection of small CC?

Evidence for dark matter in the universe

- ★ binding of galactic clusters (Zwicky, 1930s)
- ★ galactic rotation curves
- ★ large scale structure formation
- ★ inflation $\Rightarrow \Omega = \rho/\rho_c = 1$
- ★ gravitational lensing
- ★ anisotropies in cosmic MB (WMAP)
- ★ surveys of distant galaxies via SN (DE)
- ★ Big Bang nucleosynthesis
 - $\Omega_\Lambda \simeq 0.7$
 - $\Omega_{CDM} \simeq 0.25$
 - $\Omega_{baryons} \simeq 0.045$ (dark baryons ~ 0.040)
 - $\Omega_\nu \sim 0.005$



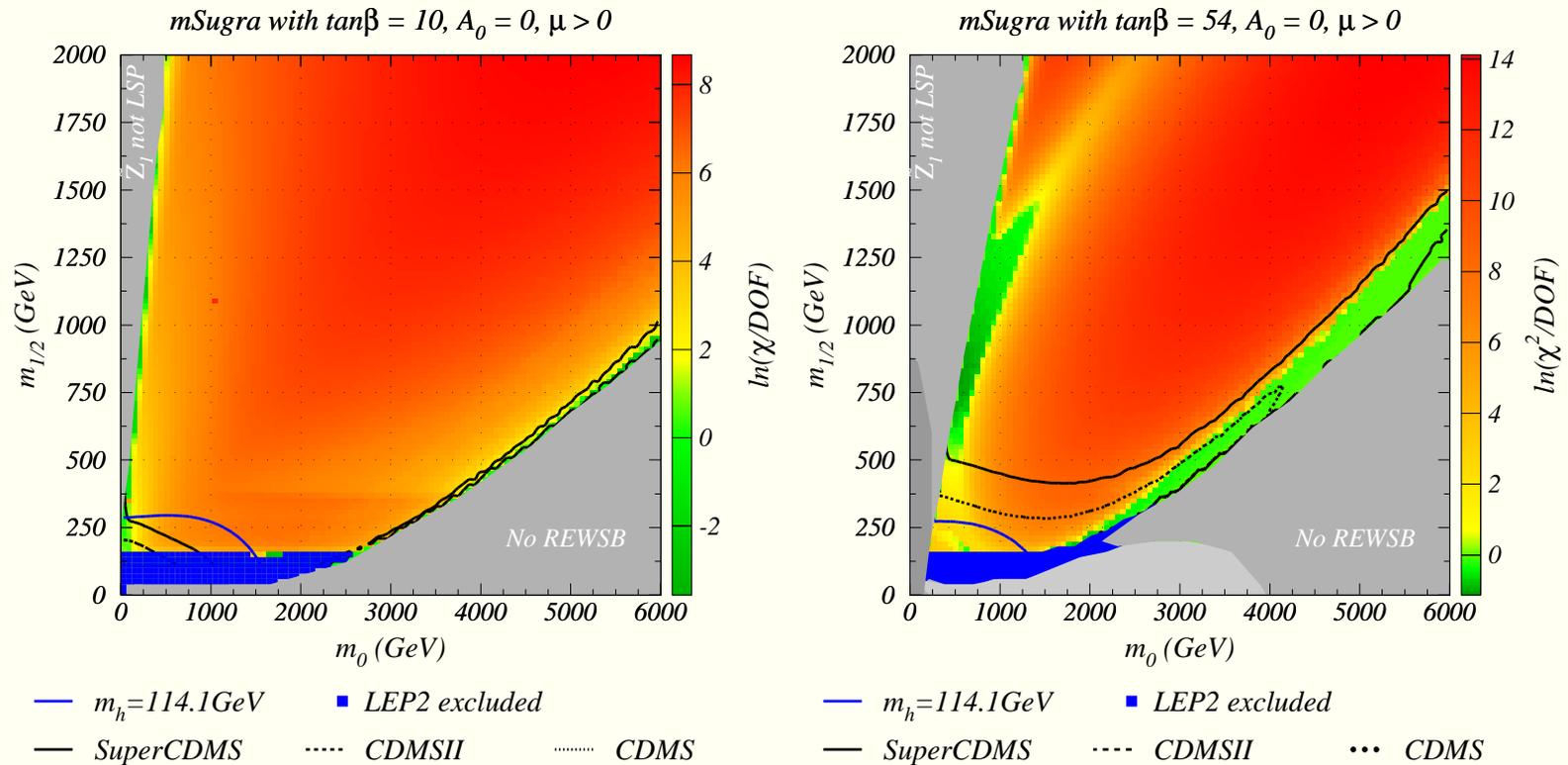
SUSY dark matter

- ★ R-parity conservation \Rightarrow conserved B and $L \Rightarrow$ proton stability
 - $R(\text{particle}) = 1; R(\text{sparticle}) = -1$
- ★ Naturally occurs in $SO(10)$ SUSY GUT theories
- ★ Some consequences:
 - Sparticles are produced in pairs
 - Sparticles decay to other sparticles
 - Lightest SUSY particle (LSP) is absolutely stable (good candidate for dark matter)
- ★ LSP must be charge, color neutral (bound on cosmological relics)
 - lightest neutralino \tilde{Z}_1 is WIMP candidate
 - axion/axino multiplet
 - gravitino

Calculating the relic density of neutralinos

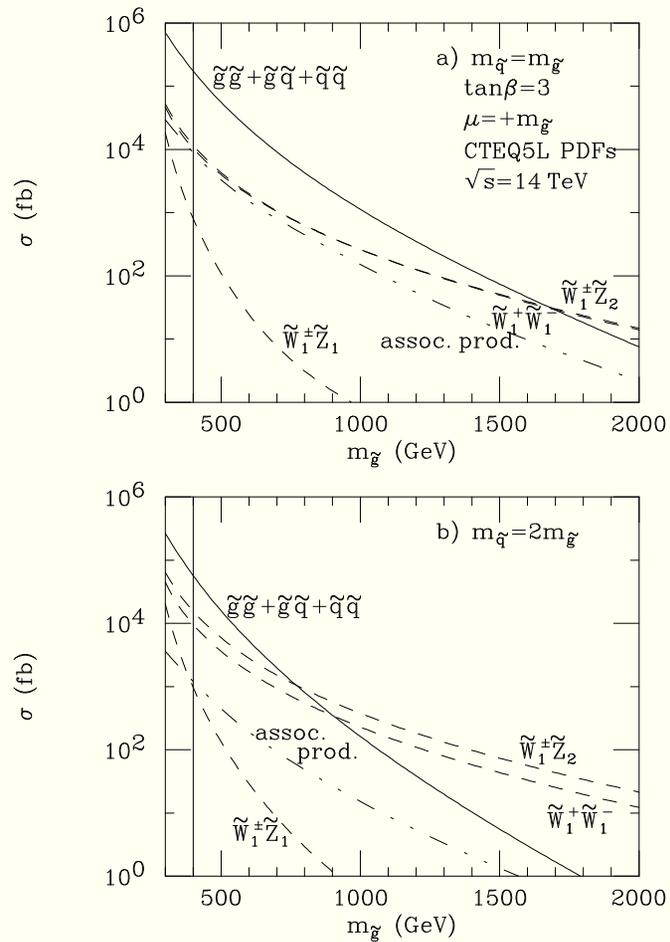
- ★ At very high T , neutralinos in thermal equilibrium with cosmic soup
- ★ As universe expands and cools, expansion rate exceeds interaction rate (freeze-out)
- ★ number density is governed by Boltzmann eq. for FRW universe
 - $dn/dt = -3Hn - \langle\sigma v_{rel}\rangle(n^2 - n_0^2)$
 - $\Omega_{\tilde{Z}_1} h^2 = \frac{s_0}{\rho_c/h^2} \left(\frac{45}{\pi g_*}\right)^{1/2} \frac{x_f}{m_{Pl}} \frac{1}{\langle\sigma v\rangle}$
 - $\Omega_{CDM} h^2 \sim 0.1 \Rightarrow \langle\sigma v\rangle \sim 0.9 \text{ pb!}$
 - $\langle\sigma v\rangle = \pi\alpha^2/8m^2 \Rightarrow m \sim 100 \text{ GeV}$
 - “The WIMP miracle!”: cosmic motivation for new physics at weak scale
- ★ SUSY: 1722 annihilation/co-annihilation reactions; 7618 Feynman diagrams
- ★ IsaReD program (HB, A. Belyaev , C. Balazs)

Results of χ^2 fit using τ data for a_μ :

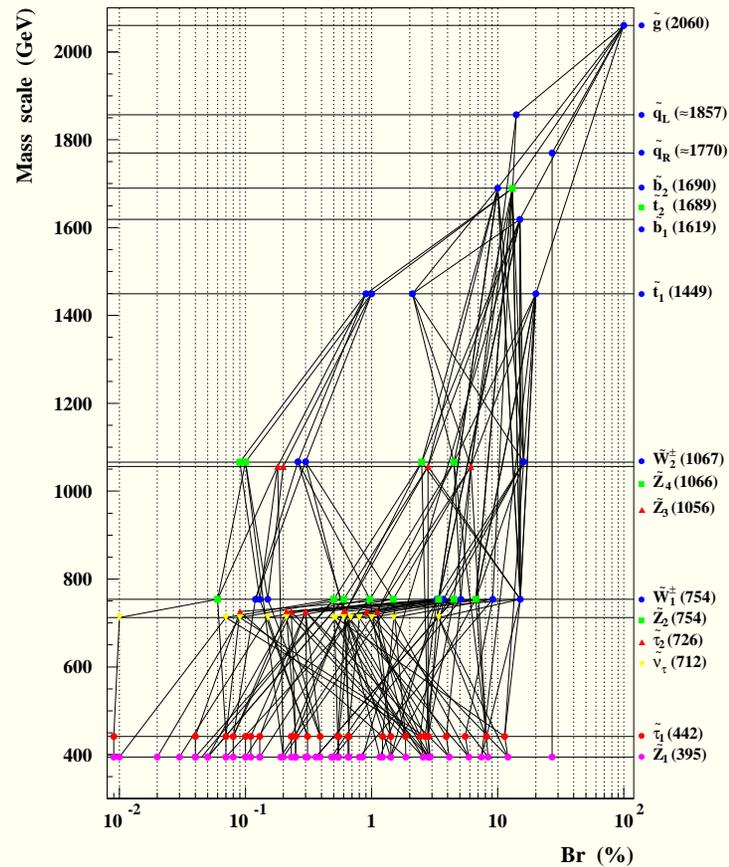


HB, C. Balazs: JCAP 0305, 006 (2003)

Production of sparticles at CERN LHC

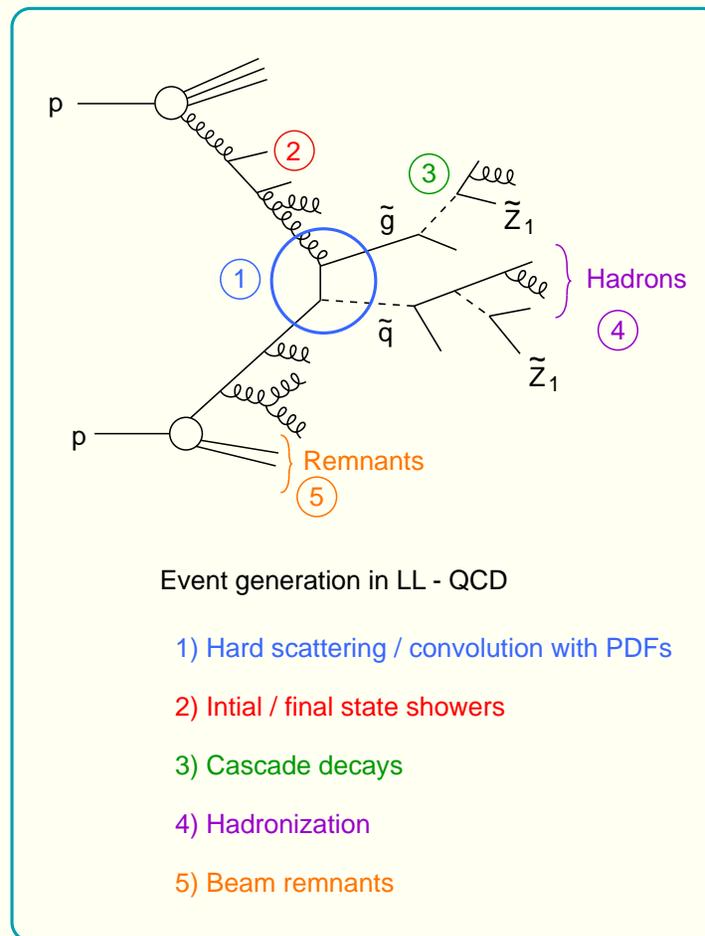


Sparticle cascade decays

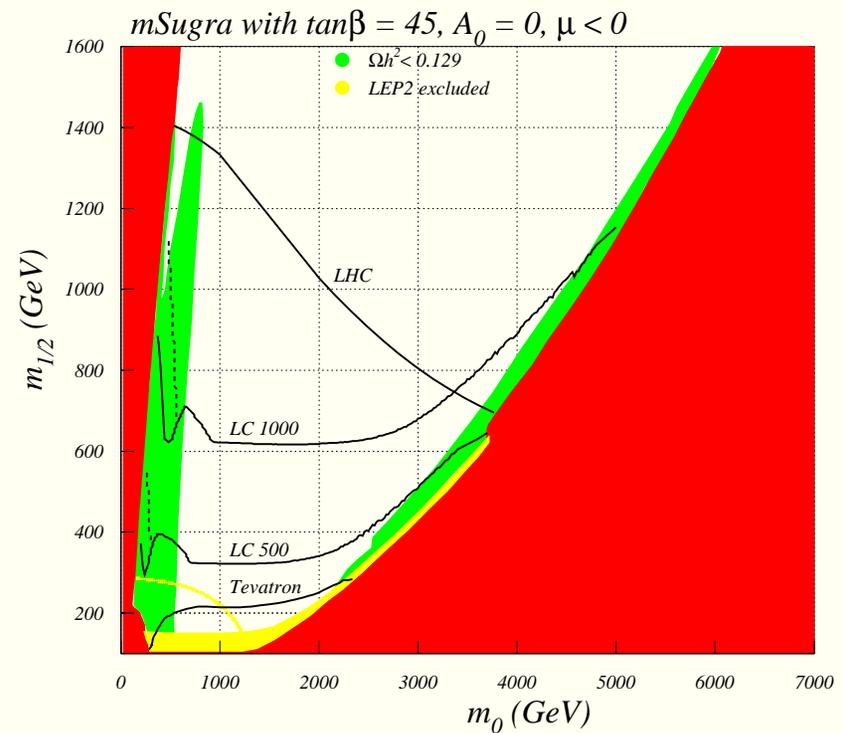
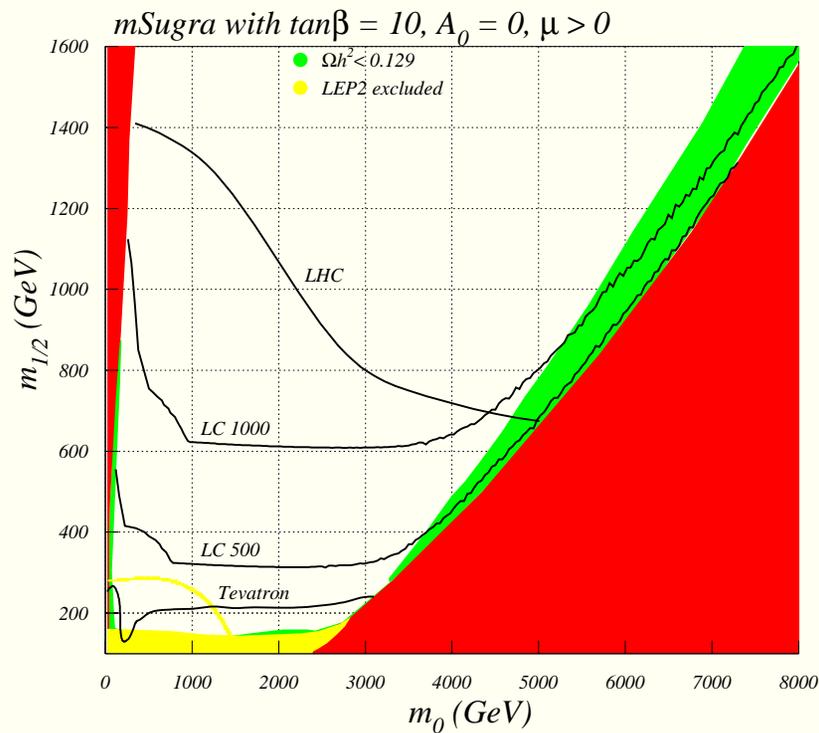


\tilde{Z}_4 qq (27.0 %)	\tilde{Z}_1 ν WWbb (4.1 %)
\tilde{Z}_1 ν Wbb (12.1 %)	\tilde{Z}_1 τ bb (2.9 %)
\tilde{Z}_1 τ WWbb (8.4 %)	\tilde{Z}_1 τ qq (2.9 %)
\tilde{Z}_1 WWbb (7.4 %)	\tilde{Z}_1 ν ZWbb (2.8 %)
\tilde{Z}_1 ν qq (5.9 %)	\tilde{Z}_1 ν hWbb (2.6 %)

Event generation for sparticles



Sparticle reach of all colliders with relic density



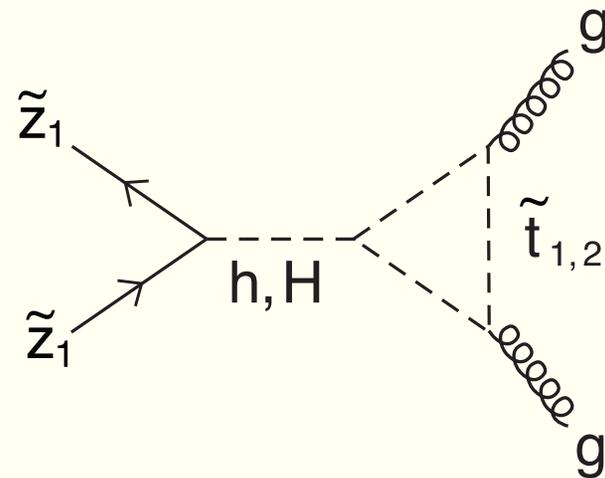
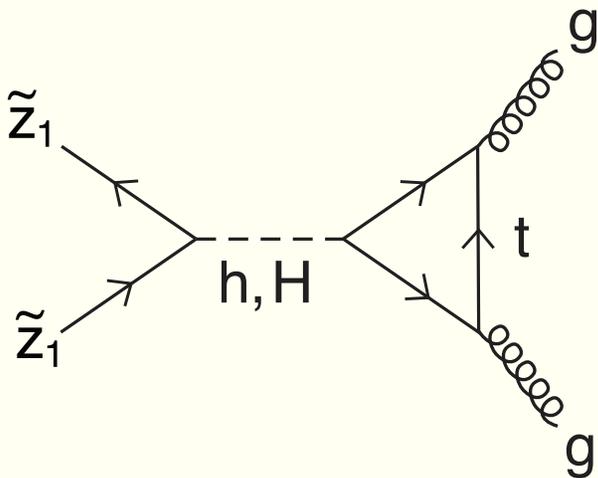
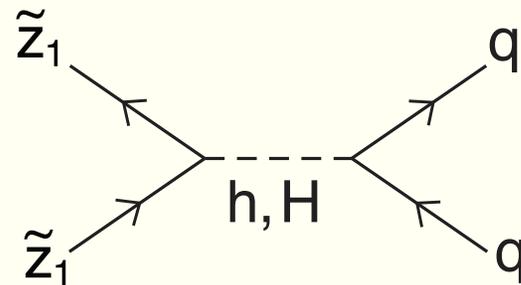
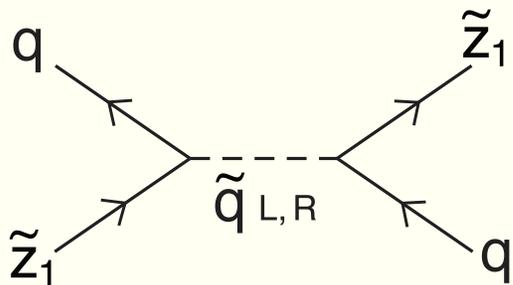
HB, Belyaev, Krupovnickas, Tata: JHEP 0402, 007 (2004)

Precision measurements at LHC: Atlas and CMS

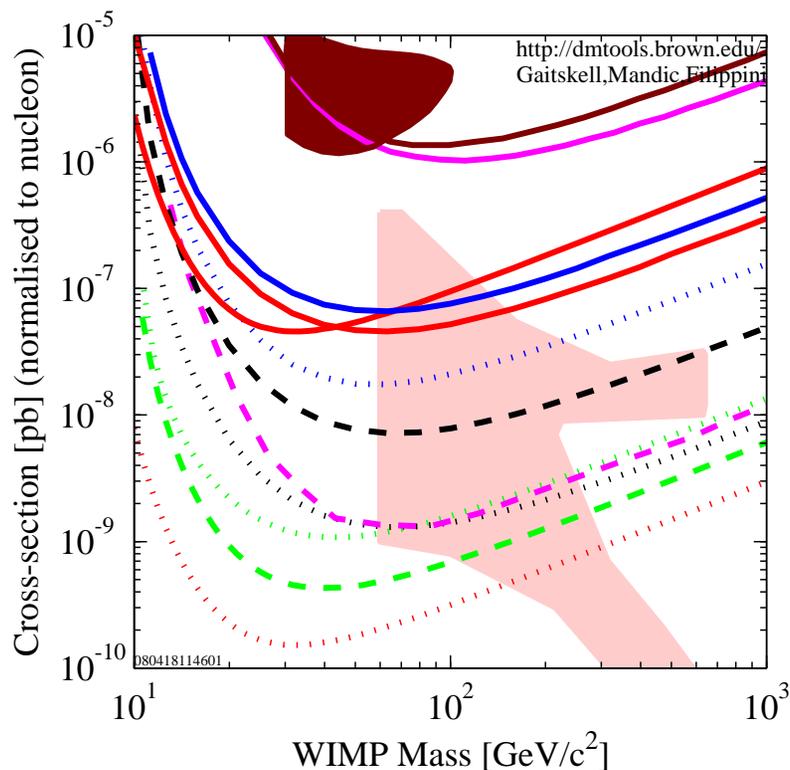
- $M_{eff} = \cancel{E}_T + E_T(j1) + \dots + E_T(j4)$ sets overall $m_{\tilde{g}}, m_{\tilde{q}}$ scale
- $m(\ell\bar{\ell}) < m_{\tilde{Z}_2} - m_{\tilde{Z}_1}$ mass edge
- $m(\ell\bar{\ell})$ distribution shape
- combine $m(\ell\bar{\ell})$ with jets to gain $m(\ell\bar{\ell}j)$ mass edge: info on $m_{\tilde{q}}$
- further mass edges possible *e.g.* $m(\ell\bar{\ell}jj)$
- Higgs mass bump $h \rightarrow b\bar{b}$ likely visible in $\cancel{E}_T + jets$ events
- in favorable cases, may overconstrain system for a given model
- ★ methodology very p-space dependent
- ★ some regions are very difficult *e.g.* HB/FP

Direct detection of SUSY DM

★ Direct search via neutralino-nucleon scattering



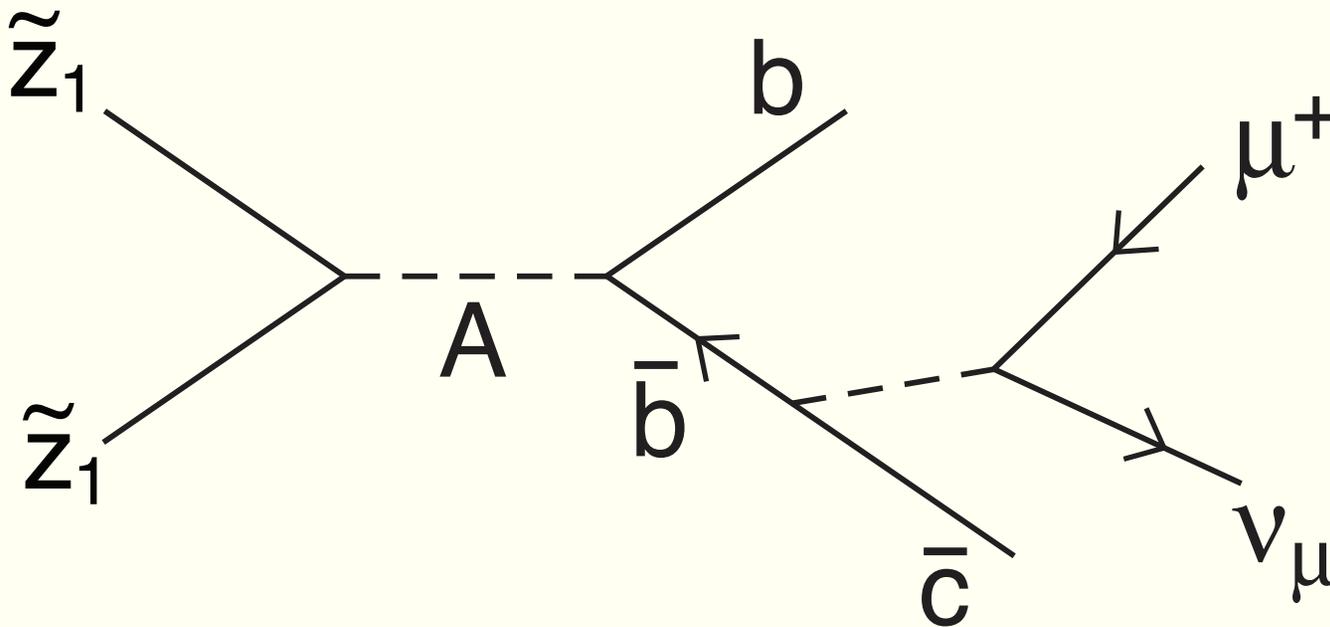
Direct detection of neutralino DM: the race is on!



- DATA listed top to bottom on plot
- Edelweiss I final limit, 62 kg-days Ge 2000+2002+2003 limit
 - DAMA 2000 58k kg-days NaI Ann. Mod. 3sigma w/DAMA 1996
 - WARP 2.3L, 96.5 kg-days 55 keV threshold
 - CDMS 2008 Ge
 - CDMS: 2004+2005 (reanalysis) +2008 Ge
 - XENON10 2007 (Net 136 kg-d)
 - CDMS Soudan 2007 projected
 - SuperCDMS (Projected) 2-ST@Soudan
 - WARP 140kg (proj)
 - SuperCDMS (Projected) 25kg (7-ST@Snolab)
 - XENON100 (150 kg) projected sensitivity
 - LUX 300 kg LXe Projection (Jul 2007)
 - XENON1T (proj)
 - Baer et. al 2003
- 080418114601

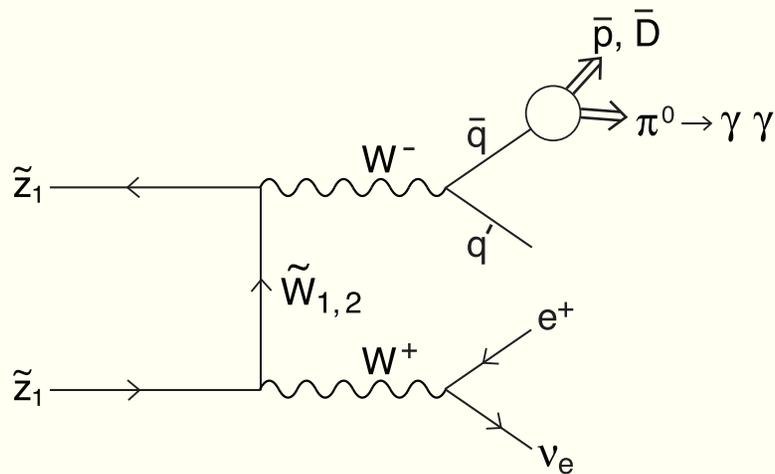
Indirect detection (ID) of SUSY DM: ν -telescopes

- ★ $\tilde{Z}_1 \tilde{Z}_1 \rightarrow b\bar{b}$, etc. in core of sun (or earth): $\Rightarrow \nu_\mu \rightarrow \mu$ in ν telescopes
 - Amanda, Icecube, Antares

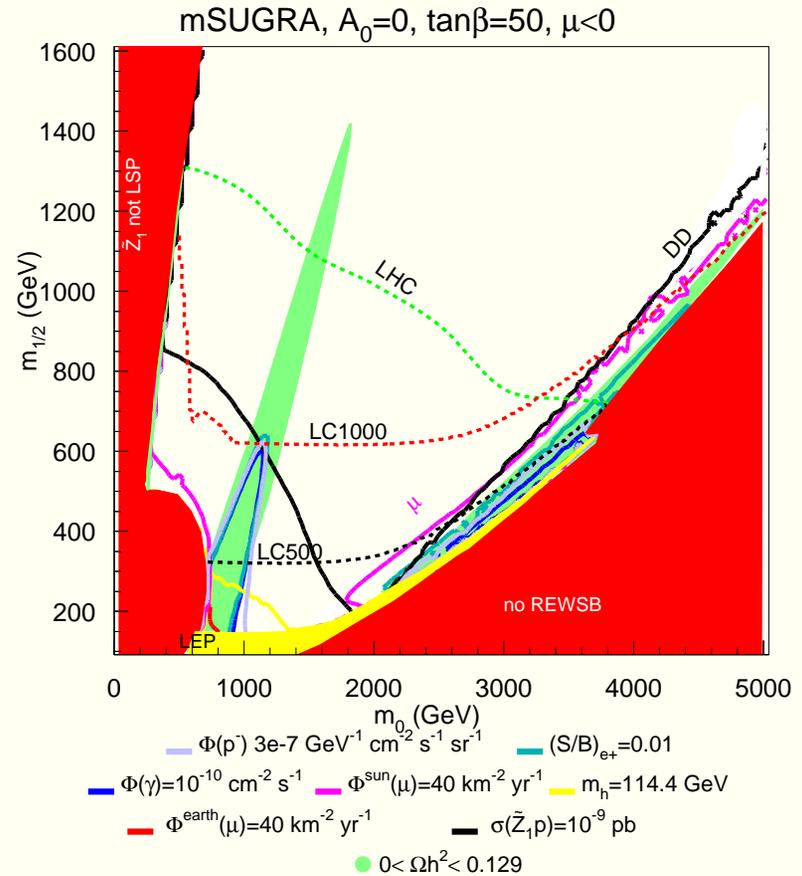
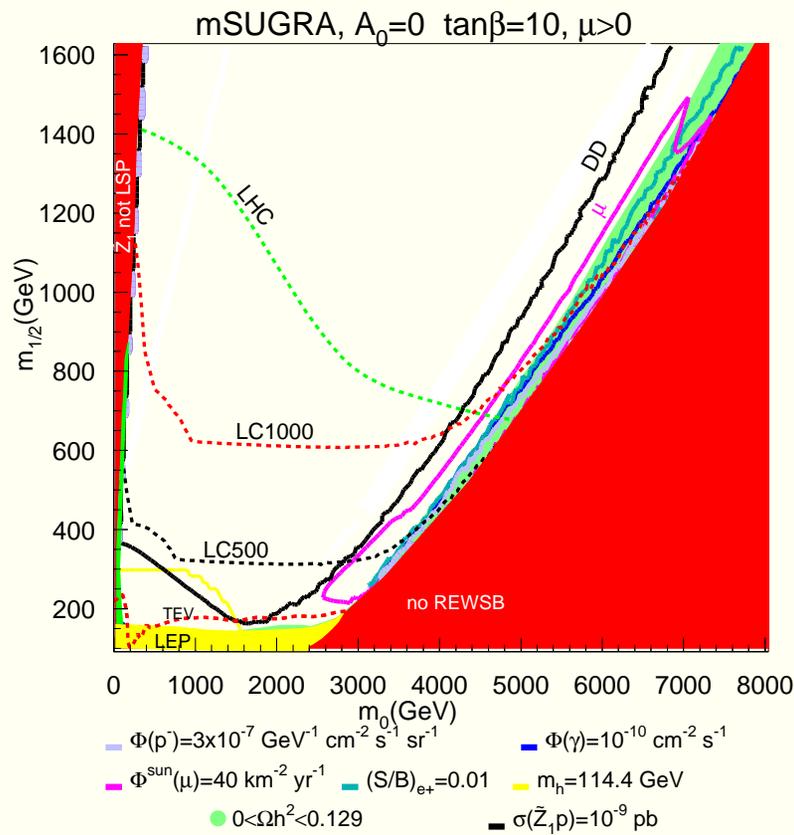


ID of SUSY DM: γ and anti-matter searches

- $\tilde{Z}_1 \tilde{Z}_1 \rightarrow q\bar{q}, \text{etc.} \rightarrow \gamma$ in galactic core or halo
- $\tilde{Z}_1 \tilde{Z}_1 \rightarrow q\bar{q}, \text{etc.} \rightarrow e^+$ in galactic halo
- $\tilde{Z}_1 \tilde{Z}_1 \rightarrow q\bar{q}, \text{etc.} \rightarrow \bar{p}$ in galactic halo
- $\tilde{Z}_1 \tilde{Z}_1 \rightarrow q\bar{q}, \text{etc.} \rightarrow \bar{D}$ in galactic halo



Direct and indirect detection of neutralino DM



HB, Belyaev, Krupovnickas, O'Farrill: JCAP 0408, 005 (2004)

Impact of DM direct/indirect detection on LHC program

- Extend reach in $\sigma_{SI} \sim 10^{-9} - 10^{-10}$ pb
 - explore thoroughly region of MHDM, possibly MWDM
- after discovery, extract m_{wimp} ?
 - $m_{\tilde{Z}_1}$ sets absolute mass scale for SUSY particles-
 - combine with LHC mass edges to gain LHC absolute sparticle masses
 - learn if \tilde{Z}_1 is absolutely stable: R -conservation
- IceCube turn-on can discover/verify especially MHDM
- knowledge of LHC spectra, σ_{SI} , σ_{SD} combined with possible gamma ray signals may allow map of dark matter distribution in the galaxy
- role of \bar{p} , e^+ , \bar{D} signals

Conclusion: We are entering a revolutionary period!

- ★ I covered just a *few* of the new physics possibilities
 - neutrino physics
 - axions and strong CP problem
 - extra dimensions
 - little Higgs models
 - supersymmetry in many guises: neutralino, axion/axino, gravitino as DM
- ★ new facilities (LHC, ν exp'ts, DM searches, ...) should lead, in the next few years, to a new paradigm— beyond the Standard Model— for the laws of physics as we know them!